History of Radio Astronomy celebrating 90 years of innovation and discovery

URSI GASS 2023,

Ronald D Ekers 23 Aug 2023, Sapporo, Japan

CSIRO SPACE & ASTRONOMY www.csiro.au



100 years of URSI 90 years of Radio Astronomy

- 1914 Brussels International Committee for Wireless Telegraphy
 Commission Internationale de Telegraphic Sans Fil Scientific
 Established to study radio communication eg fading, skip distance
- 1922 first URSI GA
 - 1929 CCIR "[radio] waves which extend over the whole earth...may perhaps even penetrate into interplanetary space
- 1933 Jansky discovers of radio emission from the centre of our galaxy
 - ➢ first presented at US URSI meeting in Washington
- 1948 URSI Commission V "Extra-terrestrial radio noise"
 ▶ 1950 renamed "Radio Astronomy"
- 1952 IAU Commission 40 "Radio astronomy" (20 years after URSI)
- 1975 URSI re-organization: commission V \rightarrow J



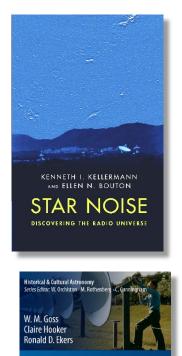
Preface to this lecture

- A historical review of discoveries in radio astronomy exploring the circumstances leading up to the discoveries
 - Many of these stories are not generally known, but they provide the background and the context.
 - I will emphasise the role played by technology
- I draw heavily on two recent publications

Kellerman and Bouton, "Star Noise" (2023) CUP Goss, Hooker and Ekers, "Joe Pawsey and the Founding of Australian Radio Astronomy" (2023) Springer

See the Springer exhibit at this meeting

- These details are often excised from the standard scientific narrative but are essential to understand the roles played by serendipity, prediction, and new technology.
- There is "nothing fortuitous" in so-called serendipitous discoveries. As Pasteur famously quoted "In the field of observation, chance favours only the prepared mind."



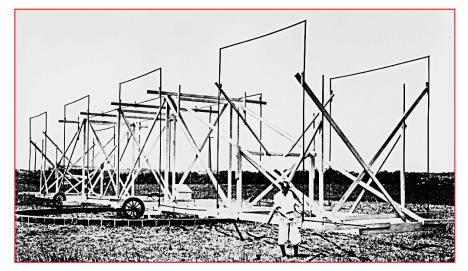
Joe Pawsey and the Founding of Australian Radio Astronomy

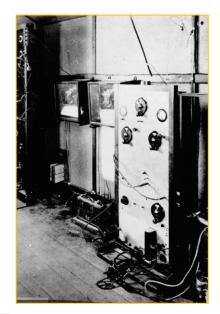
Early Discoveries, from the Sun to the Cosmos

OPEN ACCESS

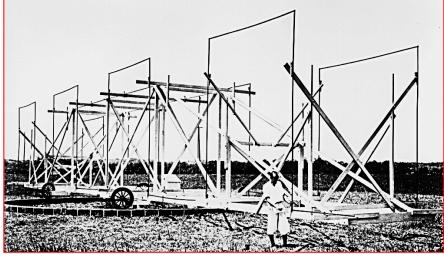
🖄 Springer











to study the effect of "atmospherics" on long distance communications.

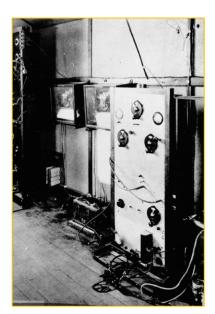
Karl Jansky, Bell Telephone

Laboratories in 1932, built this antenna

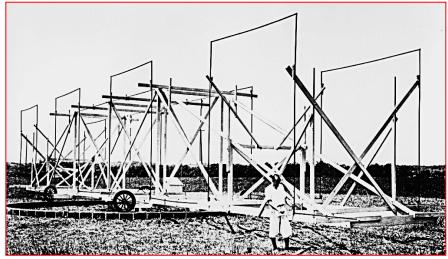
Jansky's Cosmic Hiss

Unexpected source of noise peaking each day

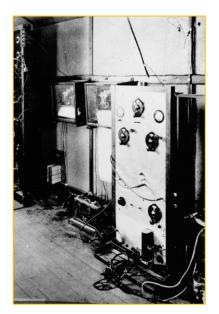
signal arrives 4 min earlier each day



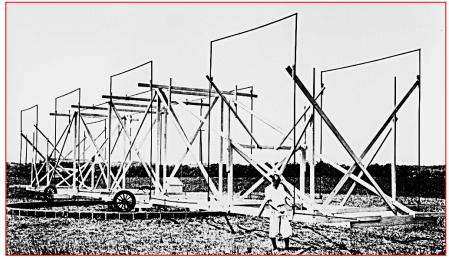




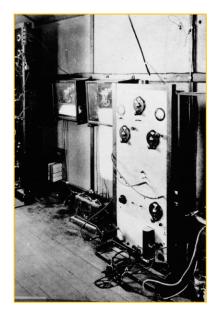
- Jansky's Cosmic Hiss
 - Unexpected source of noise peaking each day
 - signal arrives 4 min earlier each day
 - Jansky had discovered radiation coming from the center of the Milky Way.



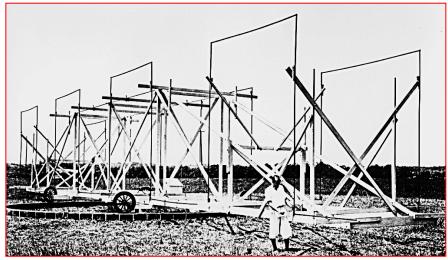




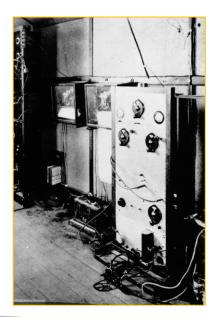
- Jansky's Cosmic Hiss
 - Unexpected source of noise peaking each day
 - signal arrives 4 min earlier each day
 - Jansky had discovered radiation coming from the center of the Milky Way.
- reaction from Bell Labs
 - "so faint not even interesting as a source of radio interference!"
 - Jansky's boss assigned him to other projects







- Jansky's Cosmic Hiss
 - Unexpected source of noise peaking each day
 - signal arrives 4 min earlier each day
 - Jansky had discovered radiation coming from the center of the Milky Way.
- reaction from Bell Labs
 - "so faint not even interesting as a source of radio interference!"
 - Jansky's boss assigned him to other projects
- not accepted by the astronomical community at the time
 - > no theoretical framework for non-thermal emission





First presentation of Jansky's discovery US National Committee for URSI 27 April 1933, Washington DC

Program.

Resolution and remarks in honor of Dr. L. W. Austin. By Dr. A. E. Kennelly (President of the U.R.S.I.), Capt. S. C. Hooper (representing the Navy Dept.), and Dr. A. H. Taylor (representing the Institute of Radio Engineers).

The effect of the electrical properties of the earth on the radiation from a simple antenna. L. P. Wheeler (Naval Research Laboratory).

A decade system for frequency comparisons. R. M. Page (Naval Research Laboratory).

Measurement of piezo-resonator constants by an impedance method. W. G. Cady (Wesleyan University).

A magnetostriction filter. Harry H. Hall (introduced by Prof. E. L. Chaffee). (Harvard University).

Audio frequency atmospherics. E. T. Burton (Bell Telephone Laboratories).

A note on hiss type atmospheric noise. K. G. Jansky (Bell Telephone Laboratories).

A compact direction finder for atmospheric disturbances. W. B. Burgess (Naval Research Laboratory).

4/27/2023

Karl Jansky

No discussion - meant nothing to anybody

URSI is "an almost defunct organization ...attended by a mere handful of old college professors."

Karl Jansky letter to his father



13



of light apparently radiated inde- of Mr. Lee, will hold himself ready pendently of starlight, originating, at all times to meet demands of the Dr. Slipher concluded, at some dis- kidnappers without police interfer tance above the earth's surface, ence.

approach him. Mr. Davis, an uncle "The provocative situation has left no alternative," the spokesman declared, deploring the fact that

facturer of Moline, Ill. Mr. Peek

has not yet been named, but his formal selection by President Roosevelt is expected as soon as the bill has been signed. As chief adhave to be so coordinated he to maintain adequate service. The President was not yet ready

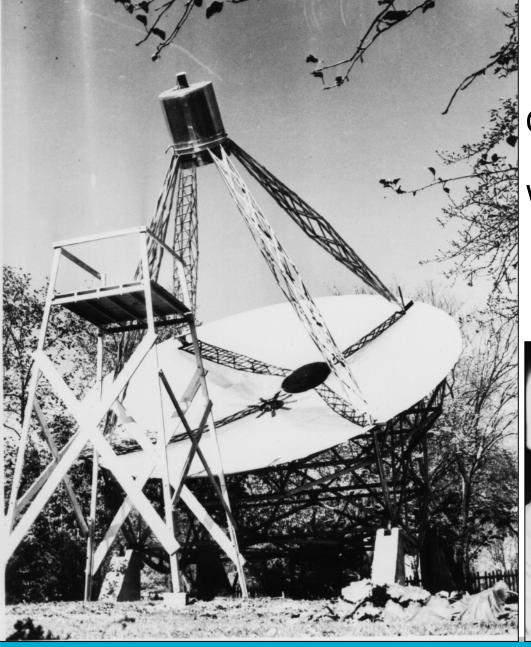
to submit to Congress "a compre-

Would Need Reservations.

Holds Positive Action Essen-Informed Roosevelt Truce

tial in Situation.

which would enable all units to work toward cooperation between capital and labor for the stabilization of industry and employment, the President's speech was received



Grote Reber

- One person did notice Jansky's discovery
- Wheaton, Illinois 1937
 - Parabolic dish for frequency flexibility
 - 32' antenna, privately funded -\$2K







Grote Reber

One person did notice Jansky's discovery

Wheaton, Illinois 1937

Parabolic dish for frequency flexibility

32' antenna, privately funded -\$2K













History of radio astronomy, URSLGASS: R D Ekers 23~Aug~2023

- 1939 finally detected cosmic static by going to longer wavelengths
 - **×** 3300 MHz
 - **×**900 MHz
 - ✓160 MHz







- 1939 finally detected cosmic static by going to longer wavelengths
 - **×** 3300 MHz
 - **×**900 MHz
 - ✓160 MHz
- Radiation had to be non-thermal
 No theoretical basis at the time
 1950 Synchrotron radiation theory
 - 10 years after Reber



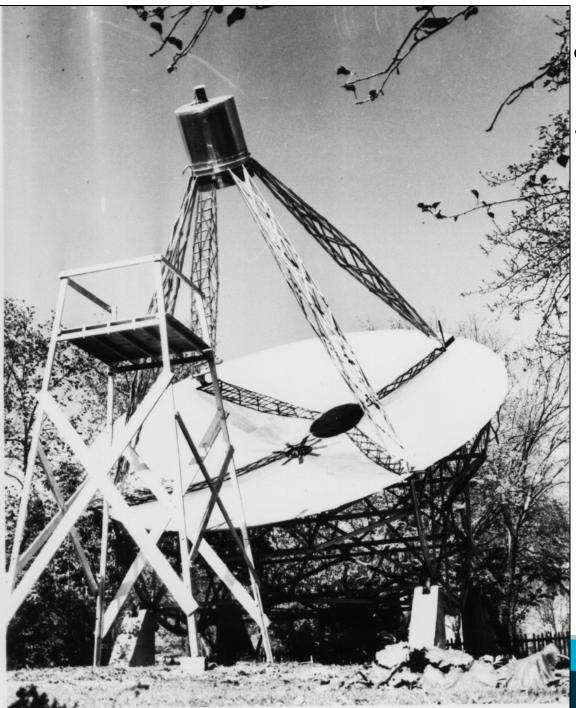


- 1939 finally detected cosmic static by going to longer wavelengths
 - **×** 3300 MHz
 - **×**900 MHz
 - ✓160 MHz
- Radiation had to be non-thermal
 No theoretical basis at the time
 1950 Synchrotron radiation theory

 10 years after Reber
- First radio map of sky
 Great difficulty getting published

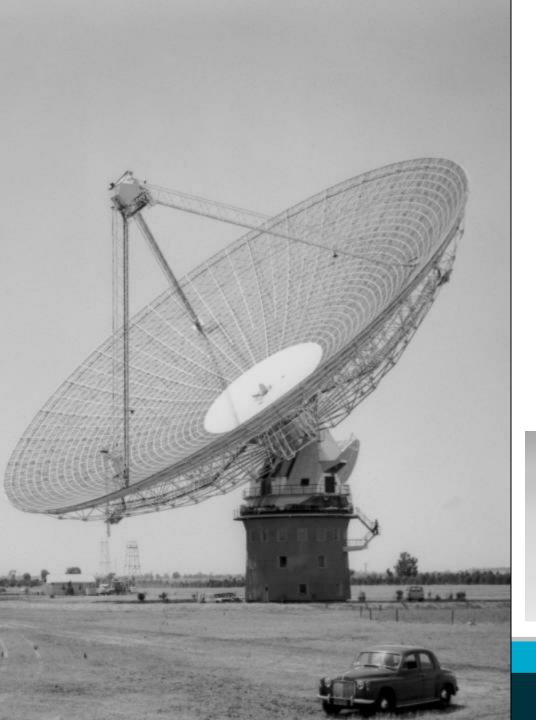






Wheaton, Illinois 1937





Parkes 64m, Australia

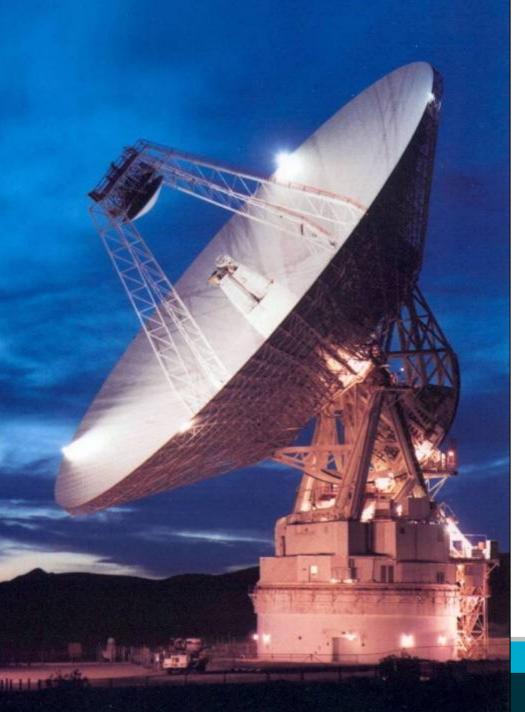
1960





Bruce Thomas hybrid mode feed





NASA Goldstone, California 1980







MPI Effelsberg 100m, 1972

Haslam 408MHz

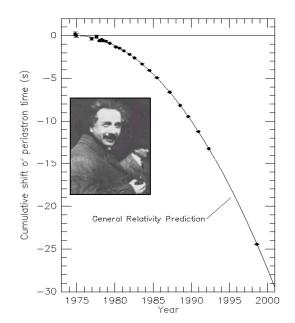


Arecibo

- 300m spherical dish
- 1963 2020
- Confirmed the predictions of General relativity
 - 1993 Noble Prize to

Taylor and Hulse

For the discovery of the binary





GBT – 2000 off axis (clear aperture) parabolic dish



SKA triggered innovation Active surface becomes locally parabolic

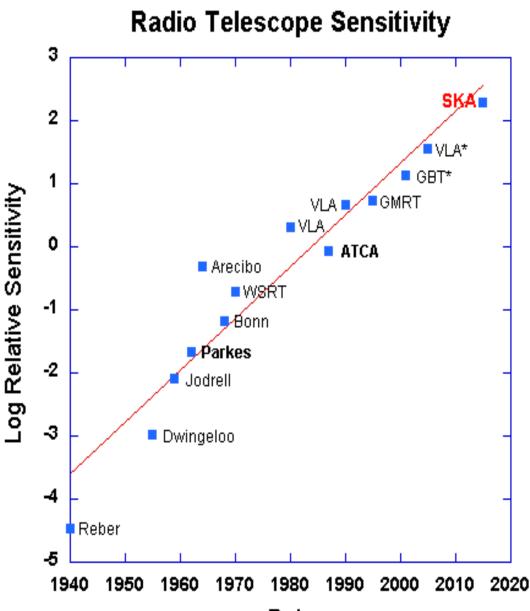


History of radio astronomy, URSI GASS: R D Ekers

FAST

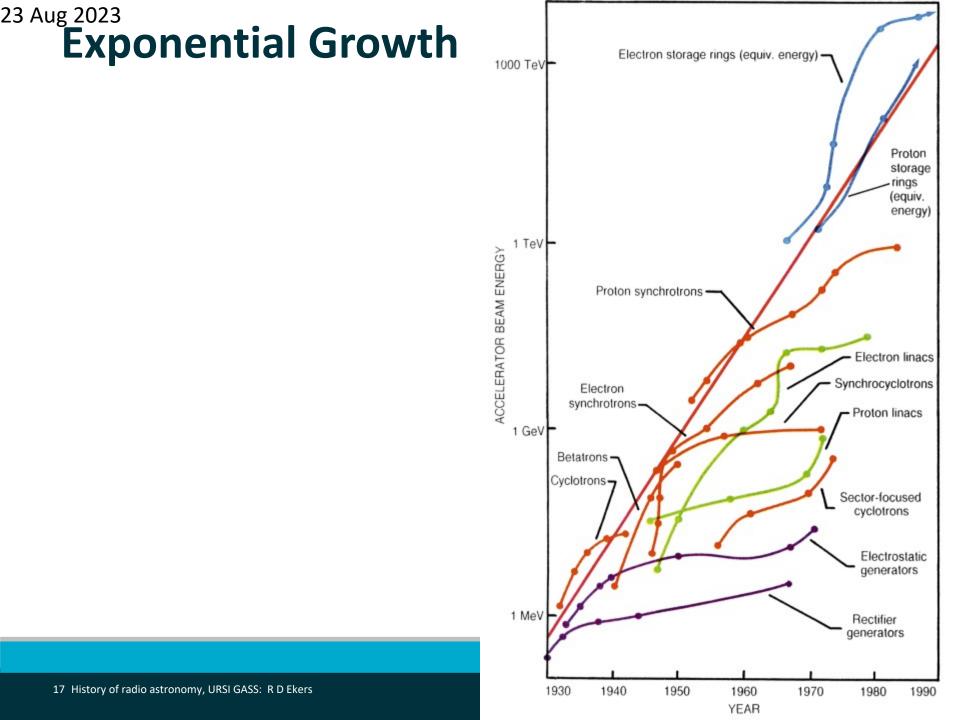
Radio Telescope Sensitivity

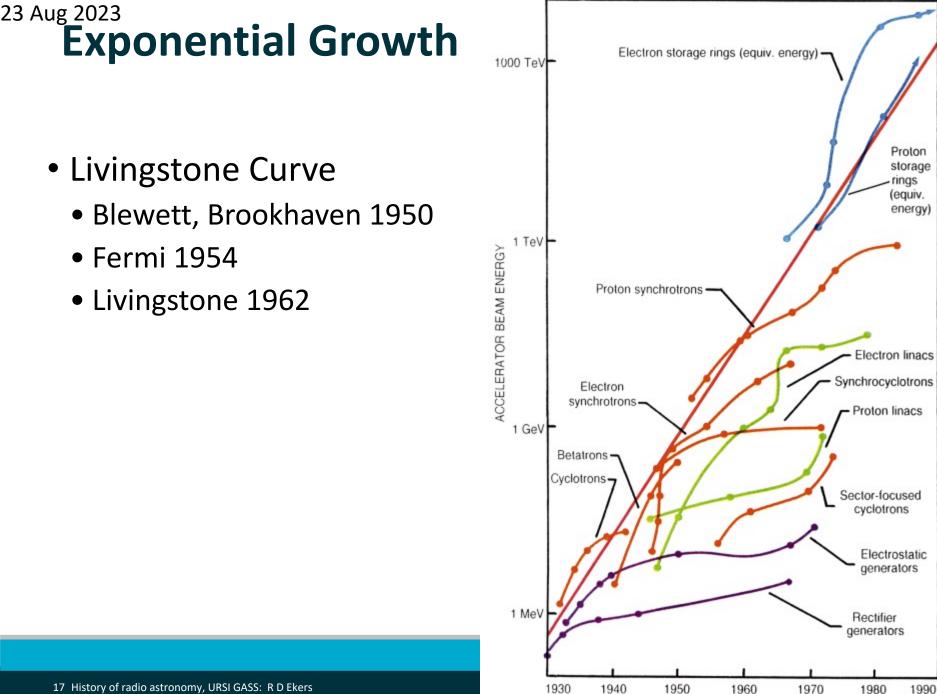
- Earlier version shown at the URSI GA Prague 1990
 General lecture
- Exponential increase in sensitivity x 10⁵ since 1940 !
 - 3 year doubling time for sensitivity
- GL- 1 Masataka Nakazawa
 Showed a similar exponential growth in communications capacity.



Date

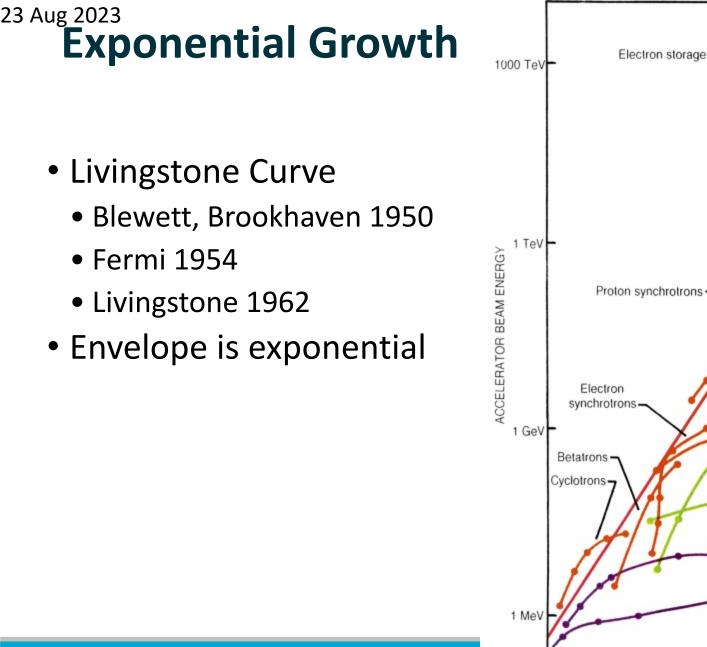


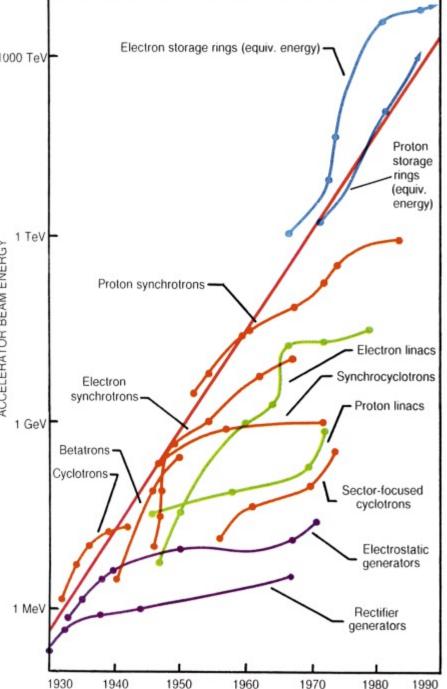




YEAR

17 History of radio astronomy, URSI GASS: R D Ekers



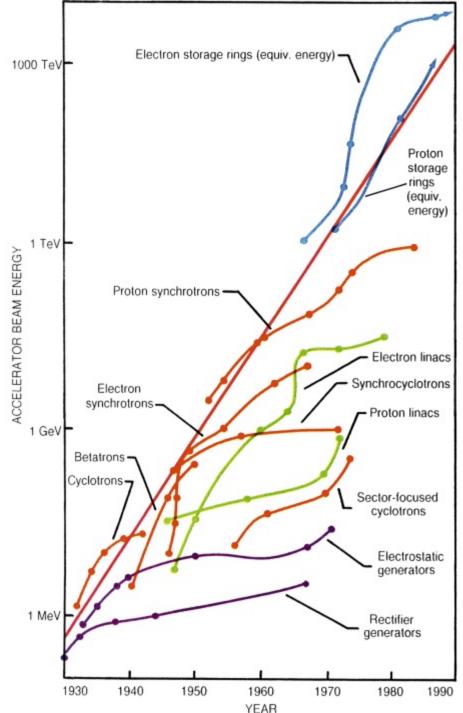


YEAR

17 History of radio astronomy, URSI GASS: R D Ekers



- Livingstone Curve
 - Blewett, Brookhaven 1950
 - Fermi 1954
 - Livingstone 1962
- Envelope is exponential
- Each technology saturates

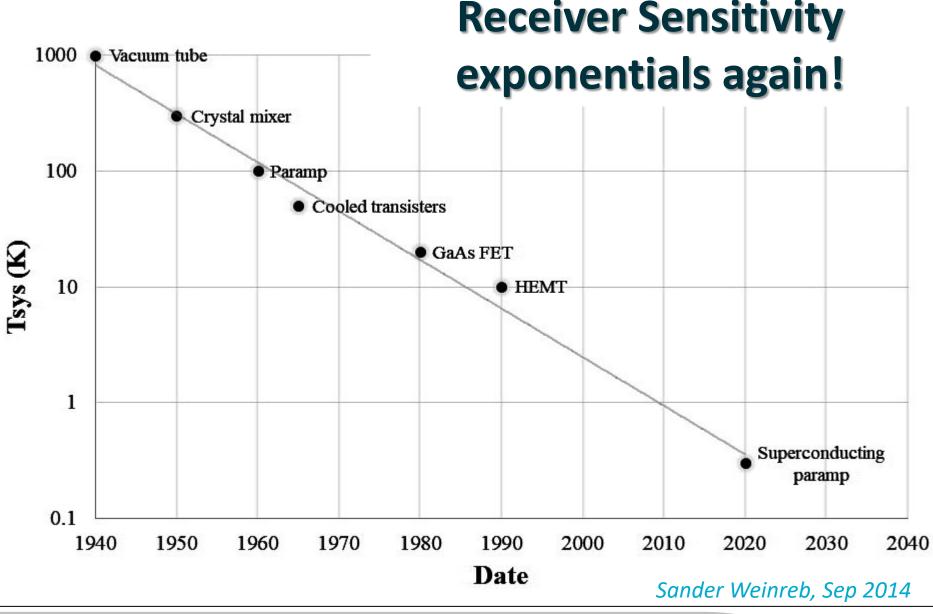


Receiver developments (Radio Astronomy)

- 1940 Vacuum tubes (>1000K)
- 1950 Crystal mixers (300K)
- 1960 Parametric amplifiers (100K)
- 1960 Masers (65K)
- 1960 Diode mixers
- 1965 Cryogenically cooled transistors (50K)
- 1980 GaAs FETs (20K)
- 1990 HEMT (10K)
- 2000 SIS (high frequency)
- 2020 Superconducting paramp (0.3K)









Technology leads scientific discoveries

De Solla Price (1963):

> most scientific advances follow laboratory experiments

Martin Harwit (1981):

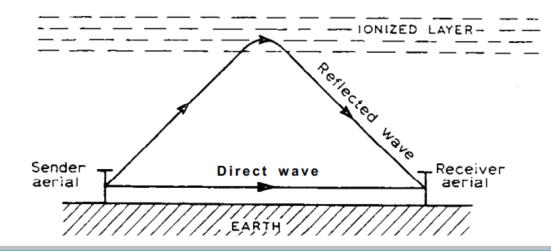
Most important discoveries result from technical innovation
 Usually within 5 years of the technical capability

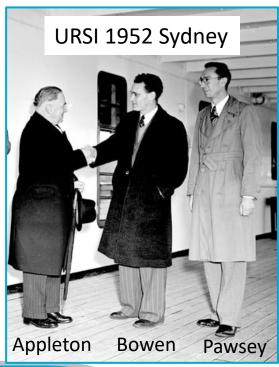
- While many discoveries are serendipitous, they depend on the development of new technology.
- It is often the telescopes, the instruments connected to the telescopes, and the data analysis that leads to new discoveries.
- The scientific discoveries for which facilities become famous are rarely those included in the initial science goals.



1924 – Edward Appleton (URSI president 1932-1952)

- Conclusive evidence for the existence of an ionosphere
 As had been proposed by Heavyside
- Measured the interference between the ground and reflected wave while sweeping frequency.
 - Separates delay and phase measured height of ionosphere
 - Appleton got the name and the Nobel prize (1974)
 - Cambridge academic physicist advantage?
- Used powerful radio broadcast transmitter



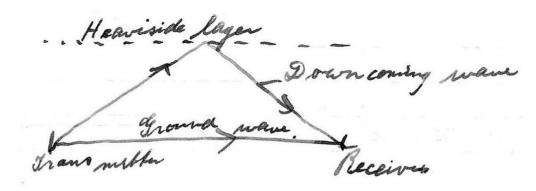




1931 – Pawsey PhD in Cambridge

- Appleton's research provided the context for Pawsey's PhD project in Cambridge with Jack Ratcliffe
- Using interference between direct and reflected radio waves to measure structure in the ionosphere







1947 – Cliff interferometer

- Pawsey wanted to measure the structure of the solar radio emission?
- Collaroy Plateau and Dover Heights near Sydney

Heaviside lager Trans mitte ecciane



Joe Pawsey and the Founding of Australian Radio Astronomy

Early Discoveries, from the Sun to the Cosmos

OPEN ACCESS

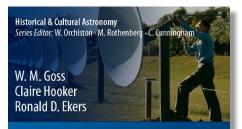
🖄 Springer



1947 – Cliff interferometer

- Pawsey wanted to measure the structure of the solar radio emission?
- Collaroy Plateau and Dover Heights near Sydney

monor



Joe Pawsey and the Founding of Australian Radio Astronomy

Early Discoveries, from the Sun to the Cosmos

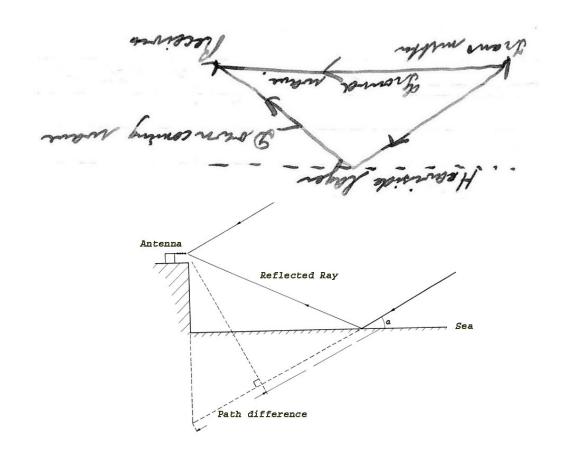
OPEN ACCESS

🖄 Springer



1947 – Cliff interferometer

- Pawsey wanted to measure the structure of the solar radio emission?
- Collaroy Plateau and Dover Heights near Sydney





Joe Pawsey and the Founding of Australian Radio Astronomy

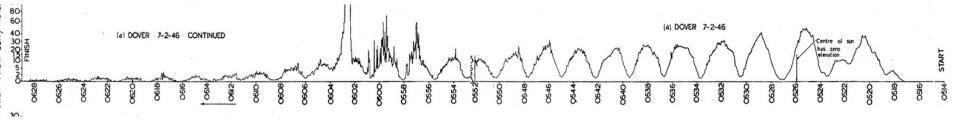
Early Discoveries, from the Sun to the Cosmos

OPEN ACCESS

🖄 Springer

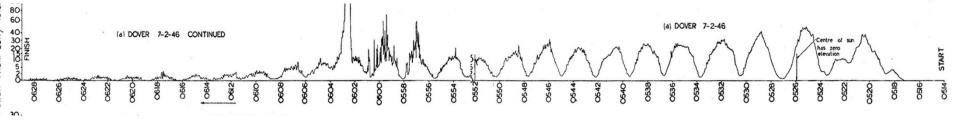


First interferometer fringes The sun, Dover Heights, 7 Feb 1946





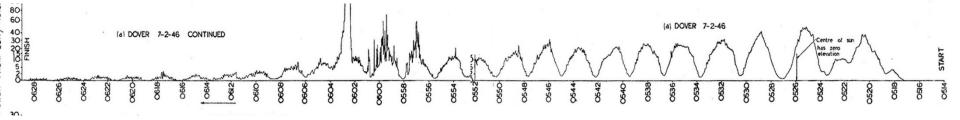
First interferometer fringes The sun, Dover Heights, 7 Feb 1946



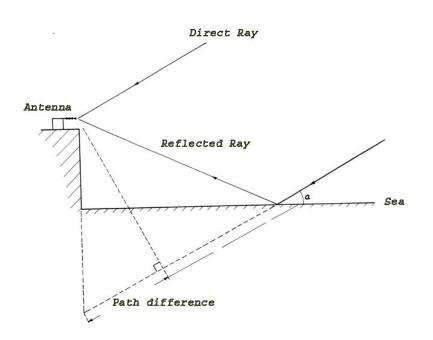




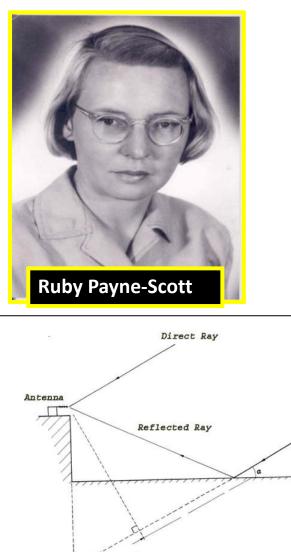
First interferometer fringes The sun, Dover Heights, 7 Feb 1946







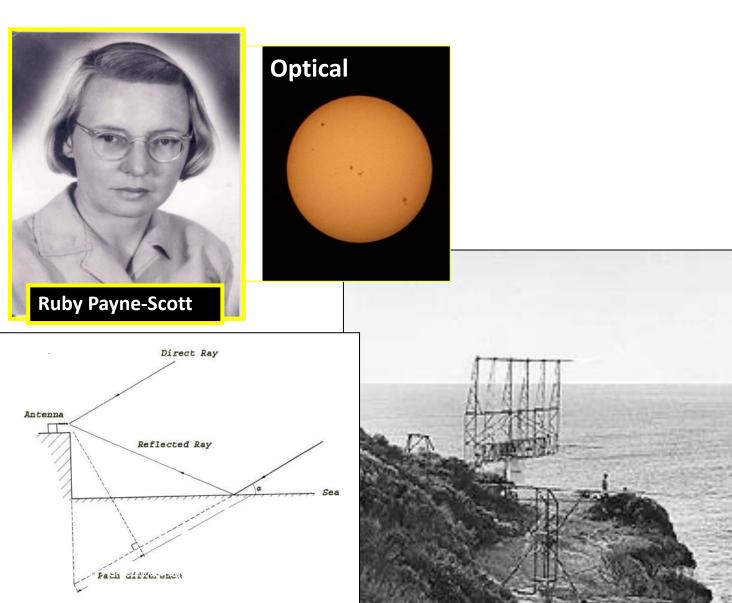




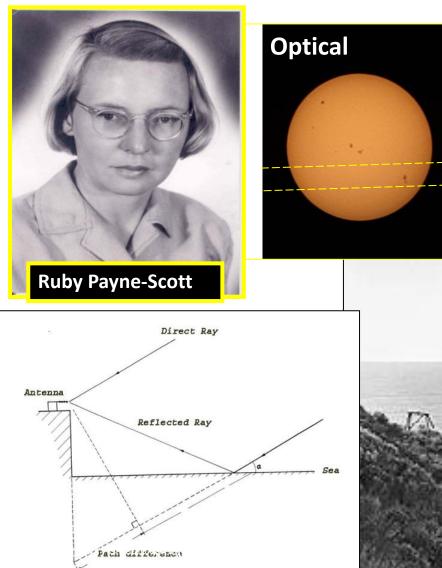
Path difference







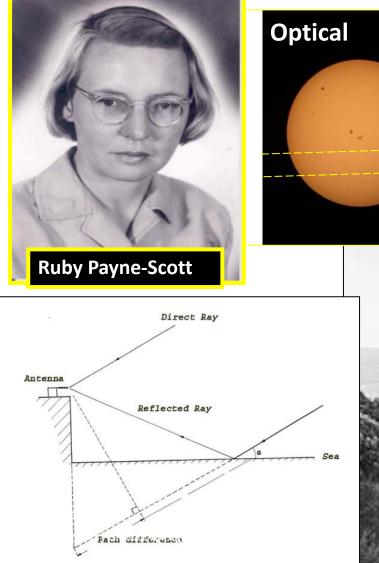




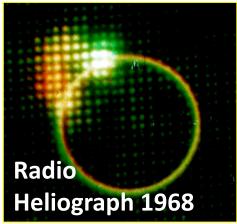
Interference pattern identifies source as sunspots

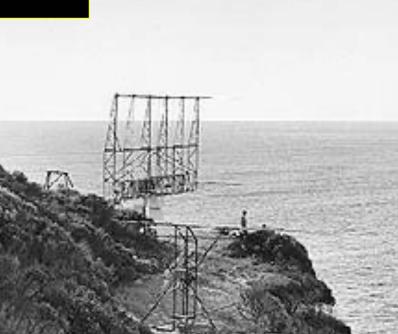




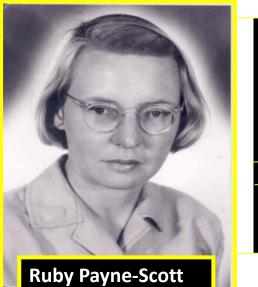


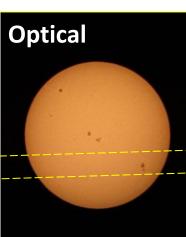
Interference pattern identifies source as sunspots



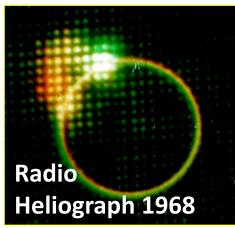


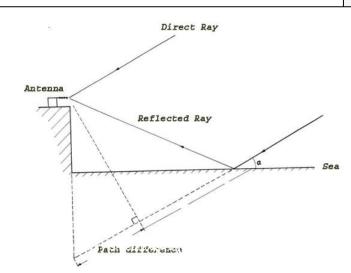




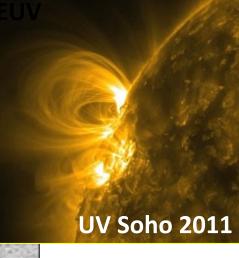


Interference pattern identifies source as sunspots







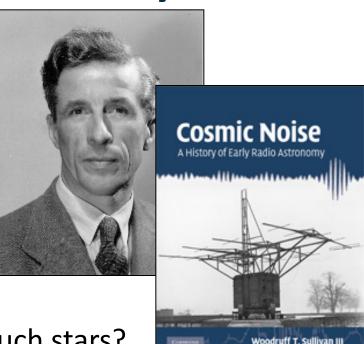




Cygnus A

strongest discrete radio source in sky

- Hey 1946
 - source with variable intensity
 time scale of seconds to minutes
 must be small diameter
 the first "radio star"
- What was it?
 - ≻no optical counterpart
 - ➤was the whole galactic plane made of such stars?





At the end of WWII the discrete radio sources were assumed to be stars

- Clever but incorrect application of Occam's razor
 - The Galaxy had been detected (Jansky)
 - ➤The sun had been detected (Hey),
 - ➤would not expect multiple different mechanisms and since the galaxy is full of stars the galactic emission must be the sum of all the stars like the sun
 - \succ Hence the discrete sources will be stars.
- Strong evidence against this interpretation was ignored
 - not enough stars to explain the galactic emission if they are like the sun - easily fixed, special bright stars
 - Evidence for extended emission ignored
 - Intensity interferometry invented to resolve the "stars"



Cliff Interferometer

- Sydney, Australia 1948
- Needed more accurate positions to identify the sources of radio emission

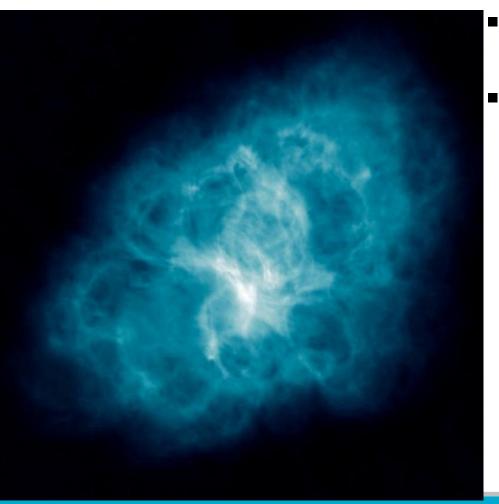




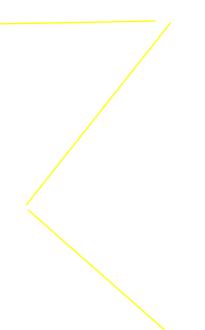


Cliff interferometer CSIRO, Australia (1948) Built to identify the radio stars (John Bolton) Discovery of the Crab Nebula radio emission Discovery of extragalactic radio sources at great distances Centaurus A Virgo A

Crab Nebula July 4, 1054 AD



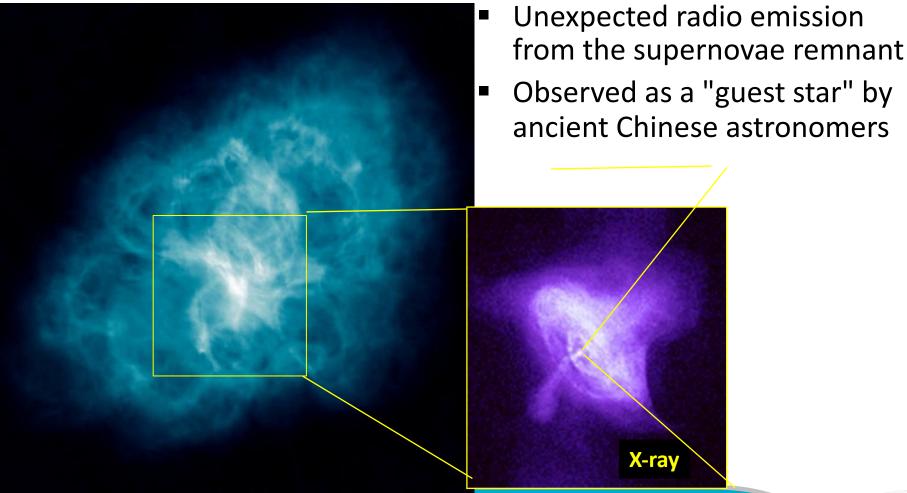
- Unexpected radio emission from the supernovae remnant
- Observed as a "guest star" by ancient Chinese astronomers





History of radio astronomy, URSLGASS: R D Ekers Radio VLA

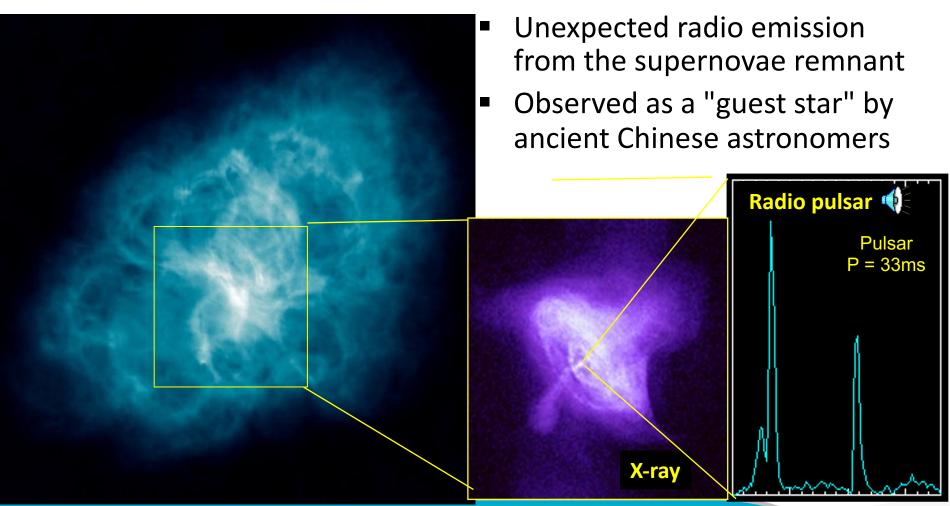
Crab Nebula July 4, 1054 AD





History of radio astronomy, UBSLGASS: R D Ekers Radio VLA

Crab Nebula July 4, 1054 AD



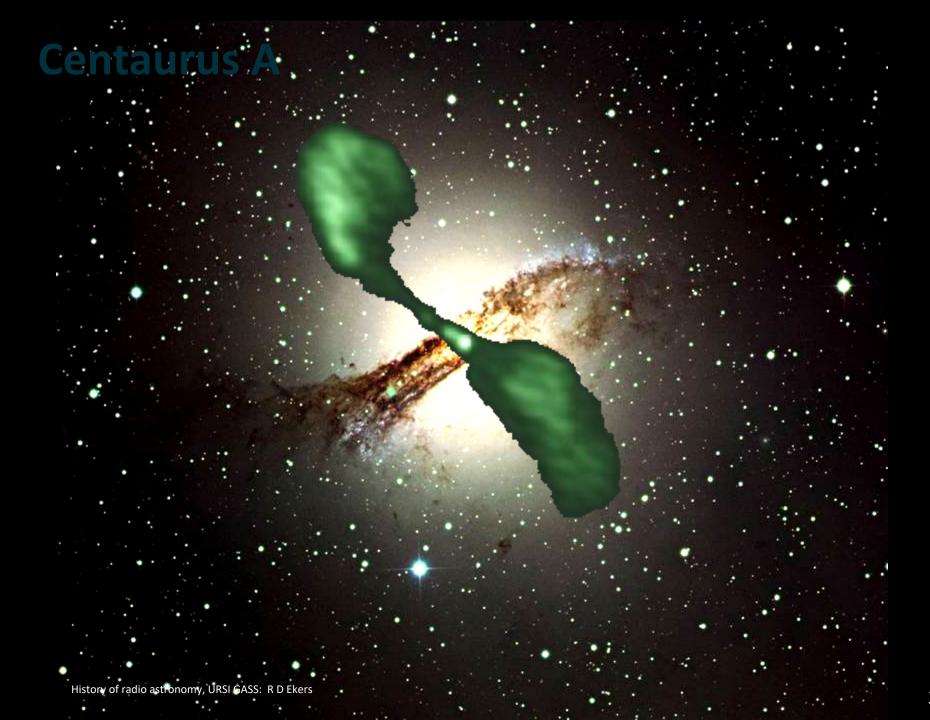


History of radio astronomy, UBSLGASS: R D Ekers Radio VLA



NGC5128 Galaxy IO in ill





Centaurus A

*

History of radio astronomy, URSLGASS: R D Ekers

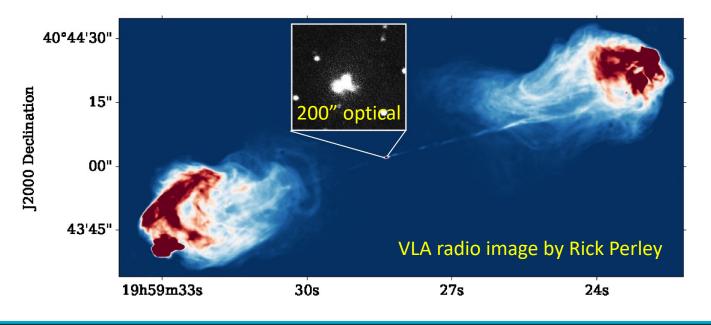


Centaurus A ATCA Mosaic



Cygnus A

- The radio galaxies were ignored until the identification of the strongest discrete radio source Cygnus A by Rudolph Minkowski
 - Colliding galaxy theory to support the identification wrong
 - But the identification was correct
 - This triggered a massive change radio galaxies at cosmological distances
 - "Cosmic noise" becomes "Radio Astronomy"





Discoveries

Kellerman and Bouton, Star Noise, 2023

- Harwit's definition
 - ➤ Named, conference,
 - Consensus (Ekers, Kellermann, Lazio, Cordes), but clearly subjective
 - Discoveries may not be a single event or a single scientist
 - > Doesn't include sub-categories. Eg radio galaxies but not tail sources....

Year	Discovery	Scientist(s)			
1933	Cosmic radio emission	Jansky ^a	1964	Interplanetary Scintillations/Solar Wind	Clarke
1938	Non-Thermal Galactic radio emission	Reber ^a	1964	Radio Recombination Lines	Dravskikh & Sorochenkc
1942	Solar corona radio emission	Reber ^a	1965	СМВ	Penzias & Wilson ^b
1943	Solar radio bursts	Неу	1965	Cosmic masers (OH)	Gunderman
1946	First Lunar Radar Detection	DeWitt, Bay	1967	Pulsars - netron stars	Bell ^c
1949	Radio galaxies	Bolton, Stanley, & Slee	1968	Water Masers	Cheung
1951	Galactic Neutral Hydrogen	Ewen	1970	CO and Interstellar molecules	Wilson, Jefferts, Penzias
1953	Double Radio Sources	Jennison & Das Gupta	1970	GR Solar Deflection	Sramek
1955	Jupiter Dekametric radio bursts	Burke & Franklin	1971	Superluminal Motion	NRAO and MIT teams
1956	Evolving Universe	Ryle	1974	Binary Pulsar	Hulse & Taylor ^b
1960	Jupiter Radiation Belts	Multiple observers ^e	1974	SgrA*	Balick & Brown
1961	, AU and Venus Rotation	, MIT & JPL teams	1979	Graviational lensing	Walsh
1962	Synthesis Imaging	Ryle ^b	1991	Exoplanets	Wolzczan, Frail
1962	Mercury Rotation& Temperature	Pettingill	1996	COBE Spectrum and Anisotropy	Mather & Smoot [⊳]
1963	Quasars	Schmidt	2007	FRBs	Lorimer
1963	First Interstellar Molecule		2019	SMBH image	Large team effort
		Weinreb, Barrett			
1964	4th Test of GR	Shapiro			



Technology enabled discoveries

Kellerman and Bouton, Star Noise, 2023

- Harwit's definition
 - Named, conference,
 - > Consensus (Ekers, Kellermann, Lazio, Cordes), but clearly subjective
 - Discoveries may not be a single event or a single scientist
 - > Doesn't include sub-categories. Eg radio galaxies but not tail sources....

Year	Discovery	Scientist(s)			
1933	Cosmic radio emission	Jansky ^a	1964	Interplanetary Scintillations/Solar Wind	Clarke
1938	Non-Thermal Galactic radio emissi	on Reber ^ª	1964	Radio Recombination Lines	Dravskikh & Sorochenkc
1942	Solar corona radio emission	Reber ^a	1965	СМВ	Penzias & Wilson ^b
1943	Solar radio bursts	Неу	1965	Cosmic masers (OH)	Gunderman
1946	First Lunar Radar Detection	DeWitt, Bay	1967	Pulsars - netron stars	Bell ^c
1949	Radio galaxies	Bolton, Stanley, & Slee	1968	Water Masers	Cheung
1951	Galactic Neutral Hydrogen	Ewen	1970	CO and Interstellar molecules	Wilson, Jefferts, Penzias
1953	Double Radio Sources	Jennison & Das Gupta	1970	GR Solar Deflection	Sramek
1955	Jupiter Dekametric radio bursts	Burke & Franklin	1971	Superluminal Motion	NRAO and MIT teams
1956	Evolving Universe	Ryle	1974	Binary Pulsar	Hulse & Taylor ^b
1960	Jupiter Radiation Belts	Multiple observers ^e	1974	SgrA*	Balick & Brown
1961	AU and Venus Rotation	MIT & JPL teams	1979	Graviational lensing	Walsh
1962	Synthesis Imaging	Ryle ^b	1991	Exoplanets	Wolzczan, Frail
1962	Mercury Rotation& Temperature	Pettingill	1996	COBE Spectrum and Anisotropy	Mather & Smoot ^b
1963	Quasars	Schmidt	2007	FRBs	Lorimer
1963	First Interstellar Molecule	Weinreb, Barrett	2019	SMBH image	Large team effort
1964	4th Test of GR	Shapiro			



Technology enabled discoveries

Kellerman and Bouton, Star Noise, 2023

- Harwit's definition
 - Named, conference,
 - > Consensus (Ekers, Kellermann, Lazio, Cordes), but clearly subjective
 - > Discoveries may not be a single event or a single scientist
 - > Doesn't include sub-categories. Eg radio galaxies but not tail sources....

Year	Discovery	Scientist(s)			
1933	Cosmic radio emission	Jansky ^a	1964	Interplanet NOBEL	nd Clarke
1938	Non-Thermal Galactic radio emission	Reber ^a	1964	Radio Reco	Dravskikh & Sorochenkc
1942	Solar corona radio emission	Reber ^a	1965	СМВ	Penzias & Wilson ^b
1943	Solar radio bursts	Неу	1965	Cosmic masers (OH)	BEL Gunderman
1946	First Lunar Radar Detection	DeWitt, Bay	1967	Pulsars - netron stars	Bell ^c
1949	Radio galaxies	Bolton, Stanley, & Slee	1968	Water Masers	Cheung
1951	Galactic Neutral Hydrogen	Ewen	1970	CO and Interstellar molecules	Wilson, Jefferts, Penzias
1953	Double Radio Sources	Jennison & Das Gupta	1970	GR Solar Deflection NOBEL	Sramek
1955	Jupiter Dekametric radio bursts	Burke & Franklin	1971	Superluminal Motion	NRAO and MIT teams
1956	Evolving Universe	Ryle	1974	Binary Pulsar	Hulse & Taylor ^b
1960	Jupiter Radiation Belts	Multiple of NORE	1	SgrA*	Balick & Brown
1961	AU and Venus Rotation	MIT & JPL NOBE		Graviational lensing	NOBEL
1962	Synthesis Imaging	Ryle ^b	1991	Exoplanets	raii ,
1962	Mercury Rotation& Temperature	Pettingill	1996	COBE Spectrum and Anisotropy	mather & Smoot [®]
1963	Quasars	Schmidt	2007	FRBs	Lorimer
1963	First Interstellar Molecule	Weinreb, Barrett	2019	SMBH image	Large team effort
1964	4th Test of GR	Shapiro			



Technology enabled discoveries

Kellerman and Bouton, Star Noise, 2023

- Harwit's definition
 - Named, conference,
 - Consensus (Ekers, Kellermann, Lazio, Cordes), but clearly subjective
 - Discoveries may not be a single event or a single scientist
 - Doesn't include sub-categories. Eg radio galaxies but not tail sources....

Year Discovery Scientist(s) Cosmic radio emission Jansky^a 1933 Interplanet ar Wind Clarke 1964 NOBEL Radio Recor 1938 Reber^a 1964 Dravskikh & Sorochenkd Non-Thermal Galactic radio emission CMB Penzias & Wilson^b 1942 Solar corona radio emission Reber^a 1965 NOBEL Cosmic masers (OH) 1943 Solar radio bursts 1965 Gunderman Hey Bell^c 1946 1967 Pulsars - netron stars First Lunar Radar Detection DeWitt, Bay Bolton, Stanley, & Slee 1968 1949 **Radio** galaxies Water Masers Cheung CO and Interstellar molecules Wilson, Jefferts, Penzias 1970 1951 Galactic Neutral Hydrogen Ewen 1970 **GR** Solar Deflection Sramek 1953 Jennison & Das Gupta **Double Radio Sources** NOBEL 1971 Superluminal Motion NRAO and MIT teams 1955 Burke & Franklin Jupiter Dekametric radio bursts Hulse & Taylor^b 1974 **Binary Pulsar** 1956 **Evolving Universe** Ryle SgrA* **Balick & Brown** 1960 **Jupiter Radiation Belts** Multiple o NOBEL Graviational lensing 1961 AU and Venus Rotation MIT & JPL NOBEL Exoplanets 1991 rail 1962 Rvle^b Synthesis Imaging mer & Smoot^b 1996 COBE Spectrum and Anisotropy 1962 Mercury Rotation& Temperature Pettingill 2007 FRBs Lorimer 1963 Schmidt Ouasars 2019 SMBH image Large team effort First Interstellar Molecule Weinreb, Barrett 1963 1964 4th Test of GR Shapiro



Nobel probability

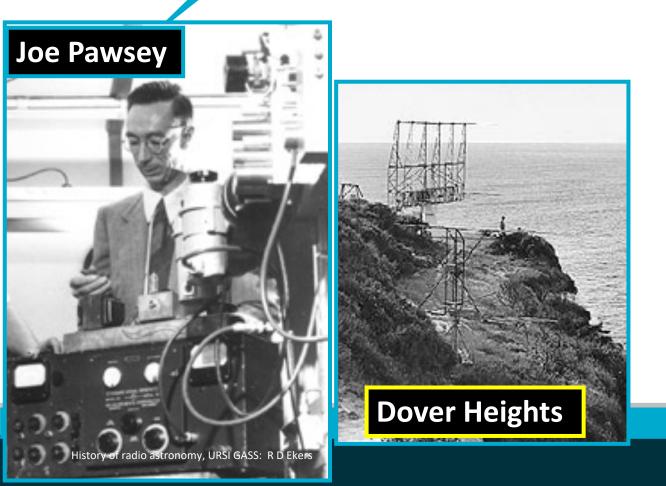
Radio astronomy

4x physics!

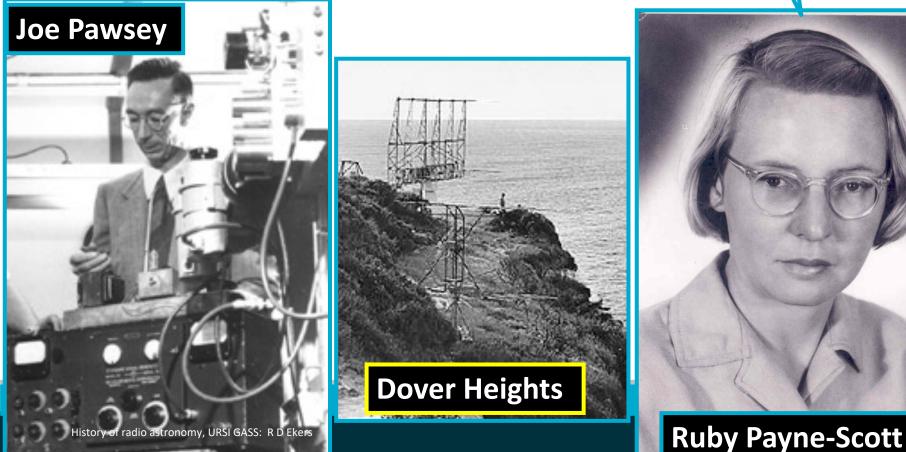
- Proc Roy Soc, Aug 1947 received July 1946!
- Used the phase of the sea interferometer fringes (lobes) to co-locate solar emission with sunspots



- Proc Roy Soc, Aug 1947 received July 1946!
- Used the phase of the sea interferometer fringes (lobes) to co-locate solar emission with sunspots



- Proc Roy Soc, Aug 1947 received July 1946!
- Used the phase of the sea interferometer fringes (lobes) to co-locate solar emission with sunspots





• It's possible in principal to determine the actual distribution by Fourier synthesis using the phase and amplitude at a range of height or wavelength.



- It's possible in principal to determine the actual distribution by Fourier synthesis using the phase and amplitude at a range of height or wavelength.
- Using wavelength as a suitable variable is unwise since the solar bursts are likely to have frequency dependent structure.



- It's possible in principal to determine the actual distribution by Fourier synthesis using the phase and amplitude at a range of height or wavelength.
- Using wavelength as a suitable variable is unwise since the solar bursts are likely to have frequency dependent structure.
- Getting a range of cliff height is clumsy, a different interference method would be more practical.





Christiansen and Warburton earth rotation synthesis (1955)





EW array makes analogue fan beams which rotate on sun





- EW array makes analogue fan beams which rotate on sun
- Chris takes the 1D FT of each strip distribution & then does a 2D Fourier synthesis using all strips
 - The way in which a 2D radio brightness distribution may be derived from a number of 1D scans is not obvious. However rather similar 2D problems have arisen in crystallography and solutions for these problems, using methods of Fourier synthesis have been found.





- EW array makes analogue fan beams which rotate on sun
- Chris takes the 1D FT of each strip distribution & then does a 2D Fourier synthesis using all strips
 - The way in which a 2D radio brightness distribution may be derived from a number of 1D scans is not obvious. However rather similar 2D problems have arisen in crystallography and solutions for these problems, using methods of Fourier synthesis have been found.
- Cites O'Brian (Cambridge) for earth rotation





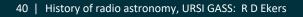
- EW array makes analogue fan beams which rotate on sun
- Chris takes the 1D FT of each strip distribution & then does a 2D Fourier synthesis using all strips
 - The way in which a 2D radio brightness distribution may be derived from a number of 1D scans is not obvious. However rather similar 2D problems have arisen in crystallography and solutions for these problems, using methods of Fourier synthesis have been found.
- Cites O'Brian (Cambridge) for earth rotation
- Swarup calculates the Fourier Transforms
 More than 1 month with electronic calculator





- EW array makes analogue fan beams which rotate on sun
- Chris takes the 1D FT of each strip distribution & then does a 2D Fourier synthesis using all strips
 - The way in which a 2D radio brightness distribution may be derived from a number of 1D scans is not obvious. However rather similar 2D problems have arisen in crystallography and solutions for these problems, using methods of Fourier synthesis have been found.
- Cites O'Brian (Cambridge) for earth rotation
- Swarup calculates the Fourier Transforms
 More than 1 month with electronic calculator







- EW array makes analogue fan beams which rotate on sun
- Chris takes the 1D FT of each strip distribution & then does a 2D Fourier synthesis using all strips
 - The way in which a 2D radio brightness distribution may be derived from a number of 1D scans is not obvious. However rather similar 2D problems have arisen in crystallography and solutions for these problems, using methods of Fourier synthesis have been found.
- Cites O'Brian (Cambridge) for earth rotation
- Swarup calculates the Fourier Transforms
 More than 1 month with electronic calculator





Computers and signal processing

- 1958
 - EDSAC II completed and applied to Fourier inversion problems
 - 360 38-point 1D transforms took 15 hours (Blyth)
 - Output was contours!





Computers and signal processing

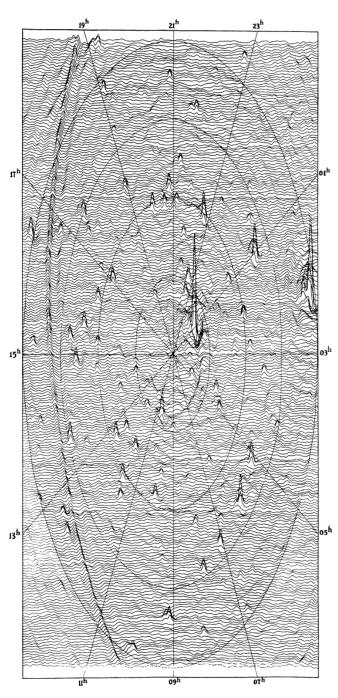
- 1958
 - EDSAC II completed and applied to Fourier inversion problems
 - 360 38-point 1D transforms took 15 hours (Blyth)
 - Output was contours!
- 1961
 - Jennison had acquired Ratcliffe's lecture notes on the Fourier transform and publishes a book on the Fourier Transform
 - Sandy Weinreb builds the first digital autocorrelator



Computers and signal processing

- 1958
 - EDSAC II completed and applied to Fourier inversion problems
 - 360 38-point 1D transforms took 15 hours (Blyth)
 - Output was contours!
- 1961
 - Jennison had acquired Ratcliffe's lecture notes on the Fourier transform and publishes a book on the Fourier Transform
 - Sandy Weinreb builds the first digital autocorrelator
- 1965
 - Cooley & Tukey publish a *convenient* implementation of the FFT algorithm



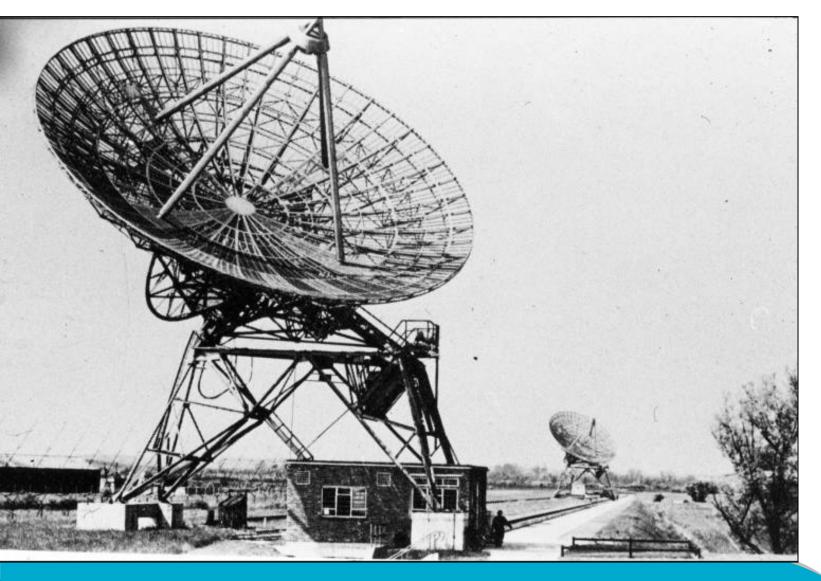


First Cambridge Earth Rotation Synthesis Image

- Ryle & Neville, MNRAS 1962
- North pole survey
- 178 MHz
- 200x200 pixels took a full night to compute on EDSACII
- Now Moore's law and the massive improvements in computing power give us LOFAR and MWA

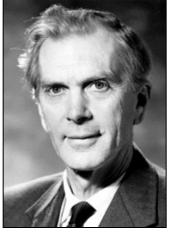


Cambridge One-Mile Telescope: 1962

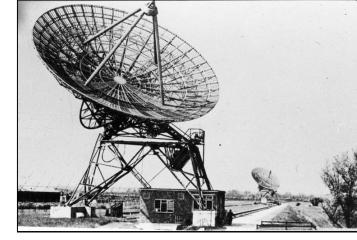




History of radio astronomy, URSI GASS: R D Ekers



Nobel Prize 1974 Sir Martin Ryle

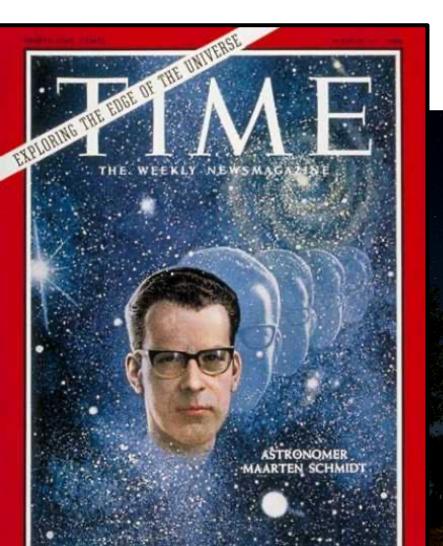


from the Nobel presentation

"The radio-astronomical instruments invented and developed by Martin Ryle, and utilized so successfully by him and his collaborators in their observations, have been one of the most important elements of the latest discoveries in Astrophysics."



Mt Palomar 200"











Hanbury Brown invents the intensity interferometer



- Hanbury Brown invents the intensity interferometer
- A few small diameter radio sources are identified with stars

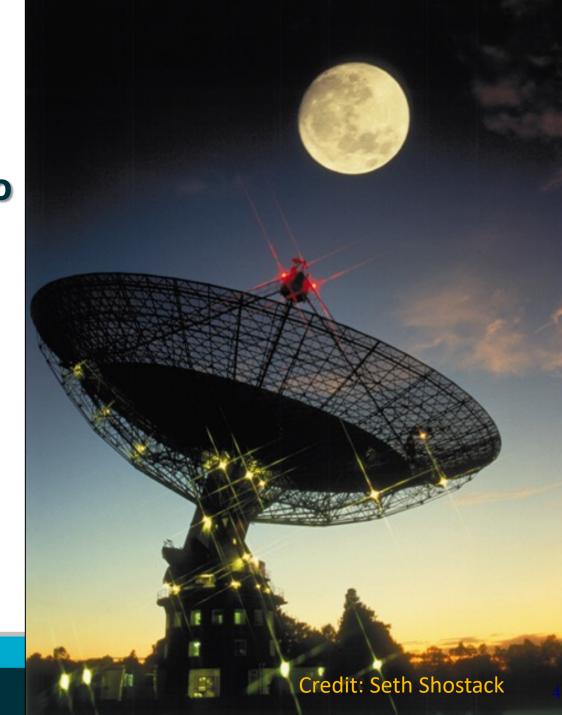


- Hanbury Brown invents the intensity interferometer
- A few small diameter radio sources are identified with stars

> These weird stars known to have variable light so "couldn't be extragalactic"

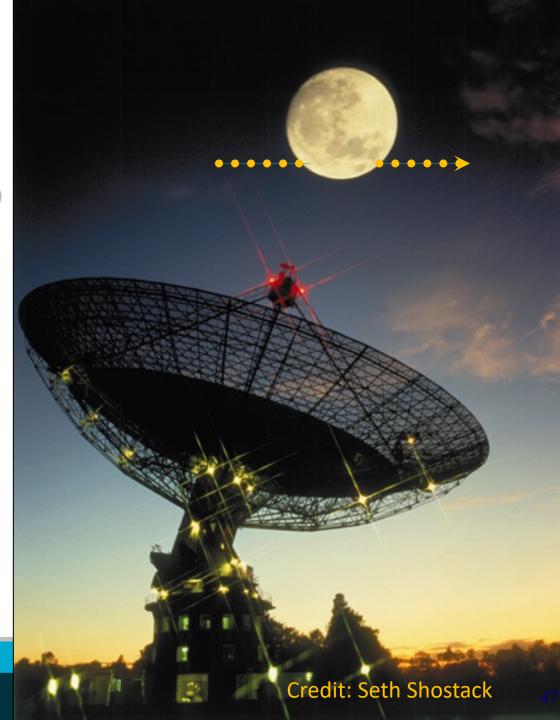


Parkes Radio Telescope observes the Lunar occultation of a radio source

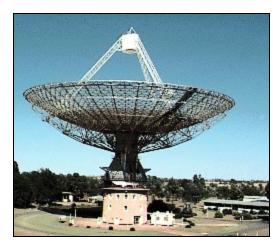


Parkes Radio Telescope observes the Lunar occultation of a radio source

- time disappearance
- and reappearance
- Core jet structure identified with a 13 mag "star" at a red shift of 0.158



Impact of the Discovery of Quasars





History of radio astronomy, URSI GASS: R D Ekers

Impact of the Discovery of Quasars

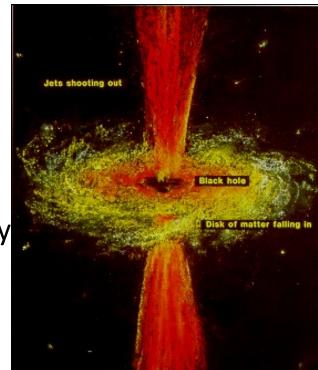
- The collision of two cultures
 Astronomers
 - General Relativity Theorists





Impact of the Discovery of Quasars

- The collision of two cultures
 Astronomers
 General Relativity Theorists
- Gravitational Collapse and Relativistic Astrophysics
 - ➢ Dallas, Texas, Dec 1963
 - Sonly gravity of a massive object in the nucleus of a galaxy could provide the energy





Steven Hawking - black holes radiate





History of radio astronomy, URSI GASS: R D Ekers

- Steven Hawking black holes radiate
- Small black holes evaporate in less than the age of the Universe





- Steven Hawking black holes radiate
- Small black holes evaporate in less than the age of the Universe
- Martin Rees a radio pulse might be observable when they disappear





- Steven Hawking black holes radiate
- Small black holes evaporate in less than the age of the Universe
- Martin Rees a radio pulse might be observable when they disappear
- John O'Sullivan and collaborators build a special instrument to look for the exploding black holes using the Westerbork radio telescopes

> There has to be a better way



IEEE 802.11 wireless network standard

- 1977 O'Sullivan explains why adaptive optics works
- 1980's Fourier Transform on a chip
 This was the *better way*
- 1990s O'Sullivan leads a multidisciplinary CSIRO team
- 1996 CSIRO obtains US patent #5,487069
- 2001 Skellern develops a wireless chip meeting IEEE standard
- 2013 1.5 billion devices sold using this technology





History of radio astronomy, URSI GASS: R D Ekers

23 Aug 2023

Westerbork: 1970



- Oort 1961 vision
- Bennelux Cross an International project
 Hogbom (Cambridge)
 +
 - Christiansen (Sydney)

⇔WSRT

- 12 x 25m dishes 1.5km
 - ➤Two moveable
 - ➤ 10 redundant spacings
 - Self calibration
- HI and dark Matter (1978)



Westerbork: 1970



- Oort 1961 vision
- Bennelux Cross an International project
 Hogbom (Cambridge)
 +

Christiansen (Sydney)

⇔WSRT

- 12 x 25m dishes 1.5km
 - ➤Two moveable
 - 10 redundant spacings
- Self calibration
- HI and dark Matter (1978)



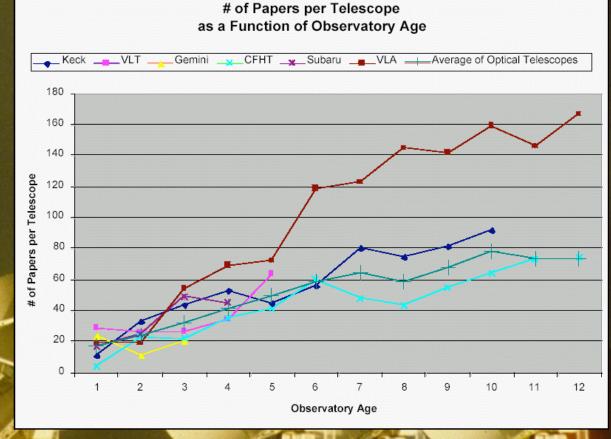
VLA New Mexico 1980

Built to observe quasars with optical resolution!



History of radio astronomy, URSI GASS: R D Ekers

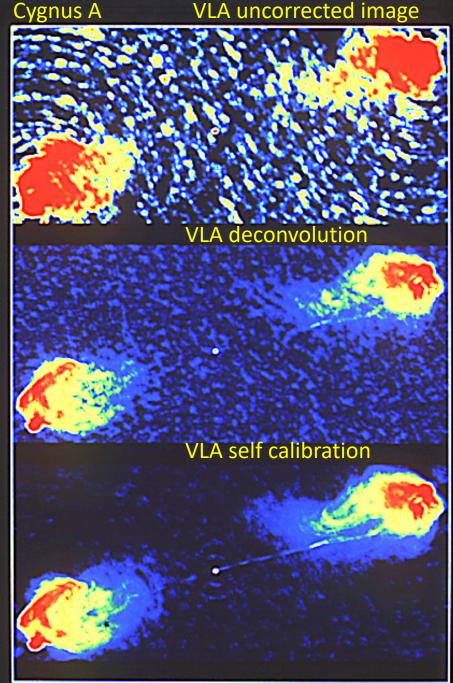
VLA New Mexico 1980



Built to observe quasars with optical resolution!



History of radio astronomy, URSI GASS: R D Ekers



Algorithms are discovered too

- VLA uncorrected image meet the design specifications
- But then we had:
 - Deconvolution
 - Self Calibration (adaptive optics)
- and now we also have
 - ➤ Mosaicing
 - Bandwidth Synthesis
 - Rotation Measure Synthesis



History of radio astronomy, URSI GASS: R D Eker

Very Long Baseline Interferometry (VLBI) from Earth and Space

- Global VLBI network
- VSOP HALCA





Very Long Baseline Interferometry (VLBI) from Earth and Space

- Global VLBI network
- VSOP HALCA



age: Blue Marble Next Generation, courtesy of Nasa Visible Earth (vi





年)春より製作・調整が行われてきた第 学術星MUSES-Bは、工学実験衛星として、大型宇宙展開アンテナ、 大容量データ伝送、高精度姿勢・軌道決定など、スペ 必要な技術試験に挑戦します。さらにこれらの技術 で宇宙の高エネルギー現象の姿を描き出す VSOP 計画の由心となります。

by Antoine de Saint-Exupéry

itute of Space and Astronautical Science



Very Long Baseline Interferometry (VLBI) from Earth and Space

- Global VLBI network
- VSOP HALCA



y Paul Boven <boven@jive.nl>. Satellite image: Blue Marble Next Generation, courtesy of Nasa Visible Earth (visibleearth.nasa.gov).



1989年(平成元年) 春より製作・調整が行われてきた第16号科 学衛星MUSES-Bは、工学実装衛星として、大型宇宙展開アンテナ、 高感度受信器、大容量データ伝送、高精度交勢・軌道決定など、スペ ースVLB観測に必要な技術試験に挑戦します。さらにこれらの技術 を総合して、超高解像度で宇宙の高エネルギー現象の姿を描き出す VSOP 計画の中心となります。

e Institute of Space and Astronautical Science



1950: 21cm Hydrogen Line Detected the standard story

- Predicted by van de Hulst
 > Leiden 1944
- Detected by Ewen & Purcell,
 > Harvard 1950
- Confirmed by Dutch and Australians



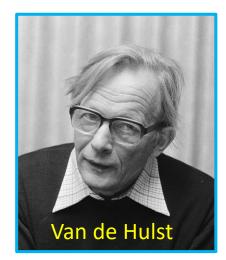




prediction \rightarrow observation \rightarrow confirmation ? but this was not the real story

Kellerman and Bouton, Star Noise (2023) CUP

- 1944 van de Hulst prediction
 - Published in Dutch and not widely accessible outside the Netherlands
- 1946 van de Hulst visits Reber in the US.
 - Reber not enthusiastic does not follow up
- 1949 Shklovskii reads review, but doesn't have v d Hulst paper
 - Re-derives the prediction from first principals!
 - Published in Russian but this is translated and distributed in the West.
 - > Landau dismisses theory no search was made in the Soviet Union
- March 1951 Atomic physicists, Ewen and Purcell, detect line
 They had read the translation of Shklovskii's paper
 - > They were interested in the atomic physics (not the astronomy)
- May 1951 Confirmed by the Dutch and the Australians

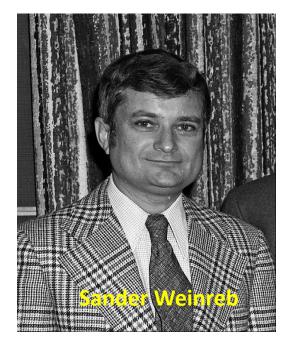




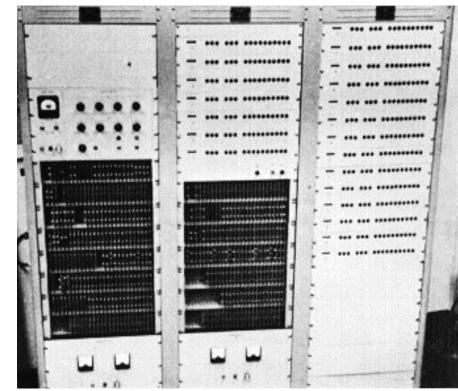


Spectral lines and the first digital correlator

- Various predictions in 1955 including OH, NH₃, H₂O
- OH detected by Weinreb and Barret 1963
 - > digital auto-correlation spectrometer
 - ➢ Parkes confirmed in only 1 day with a receiver modification



21 lags 300kHz clock discrete transistors \$19,000





There were many more spectral lines Massive clouds of Molecules in the space between the stars

Radio Recombination Lines (RRL)

Soviet prediction and detection in 1964 but not believed in West.

- Theoreticians repeatedly discouraged observers incorrect estimate of Stark broadening
- > These lines were easily detected with the new 140' telescope in 1965
- OH line detected in 1965 but initially called mysterium
- NH₃ 1967
- H₂O 1968
 - detected at Berkeley but NRAO had rejected proposal to look
 - \succ H₂O was so strong it could have been detected before HI!
- By 1969 this was a thriving research area with many more lines being detected



The Masers – more serendipity

- Theorists were aware of the maser amplification process but the required conditions in the interstellar medium were considered so unlikely that this was never pursued
 - Maser amplification first recognised through observations of spectral lines of OH with 100% polarization
 - Maser amplification was confirmed by high angular resolution (VLBI) observations

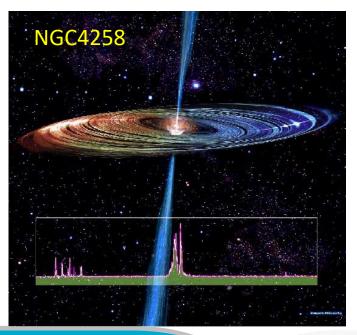


Extragalactic Masers

- Extragalactic H₂O mega masers were discovered in 1971
- NGC4258 H₂O maser detected at Nobeyama in 1992
 - This provided the first direct evidence for a black hole in the nucleus of a galaxy
 - broad (1000km/s) line and VLBI size -> 36 x 10⁶ solar mass black hole

Myoshi et al (1995) [SN#77]

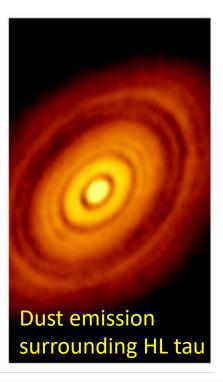






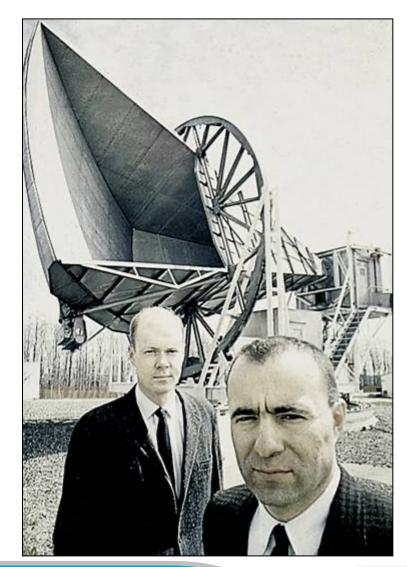
ALMA and mm radio astronomy

- CO was not detected until 1970
 > receivers at 2.6mm were too difficult before then
- A huge new area of astronomy has now opened up but I have no time to include this in my review



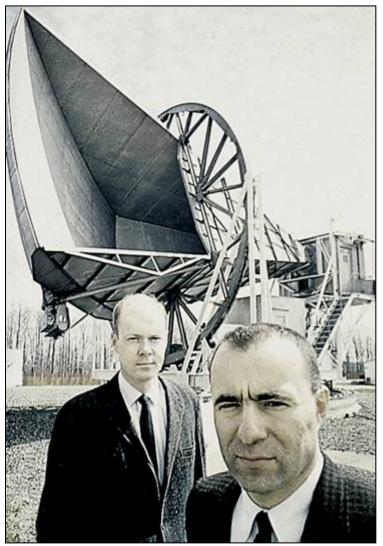






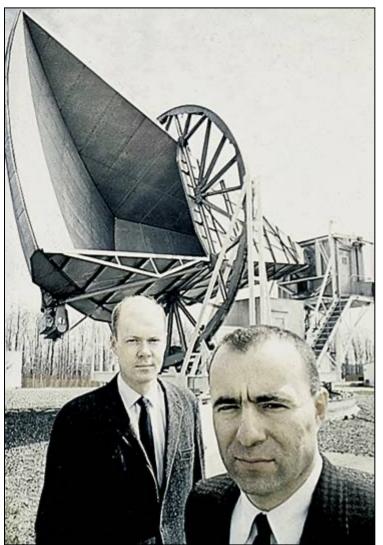


 Two radio astronomers working on the new satellite communications systems at Bell Labs discovered radio waves from the cosmic fireball at the beginning of the universe



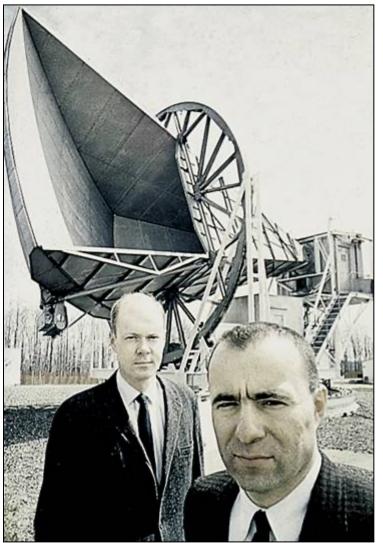


- Two radio astronomers working on the new satellite communications systems at Bell Labs discovered radio waves from the cosmic fireball at the beginning of the universe
- 1978 Nobel prize to Penzias and Wilson for discovery of the Big Bang radiation

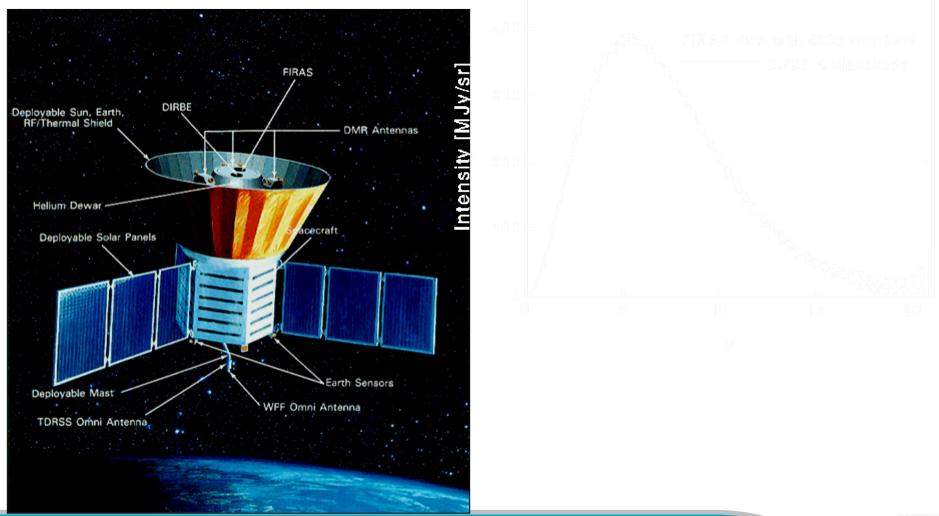




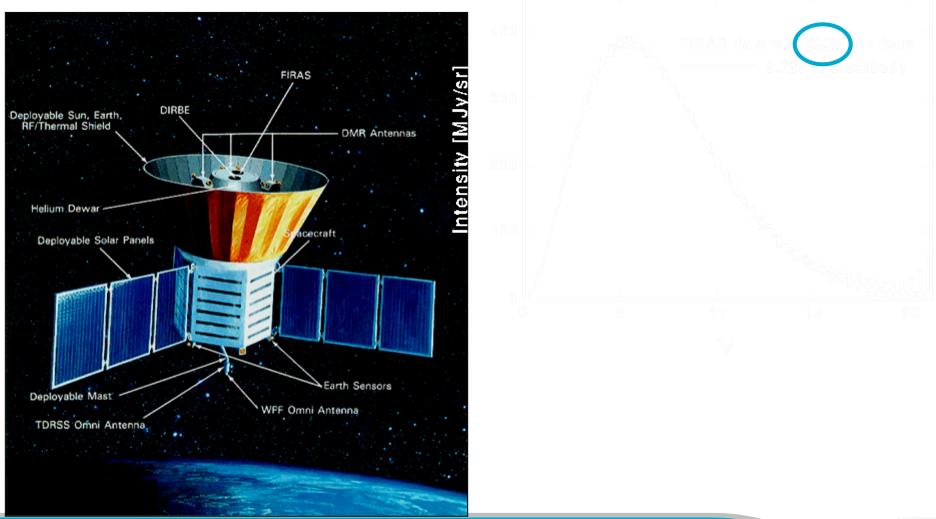
- Two radio astronomers working on the new satellite communications systems at Bell Labs discovered radio waves from the cosmic fireball at the beginning of the universe
- 1978 Nobel prize to Penzias and Wilson for discovery of the Big Bang radiation
- Serendipitous observation of a predicted phenomena
 - Bob Dicke's experiment to search for this was already in progress at Princeton



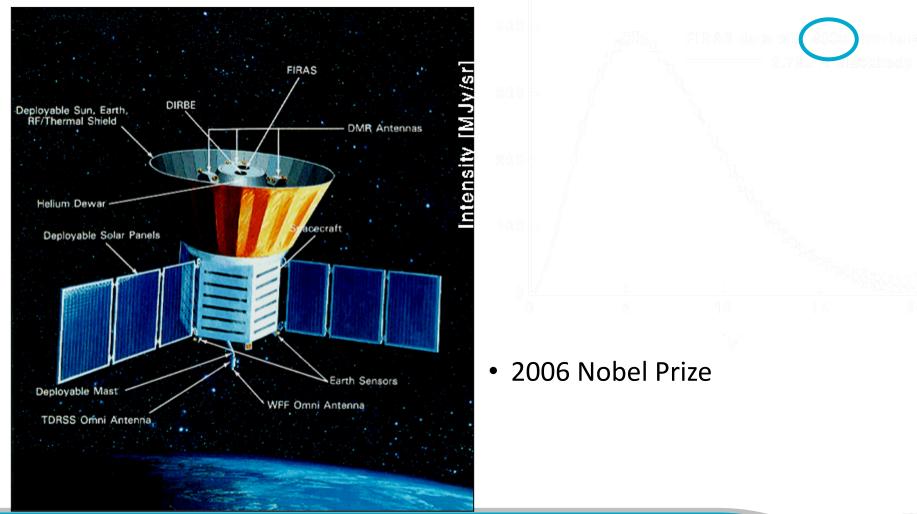






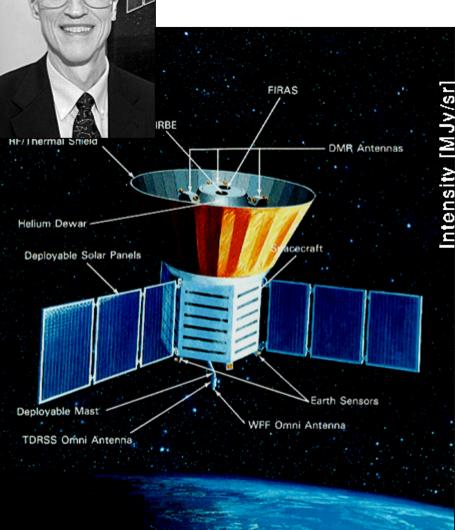


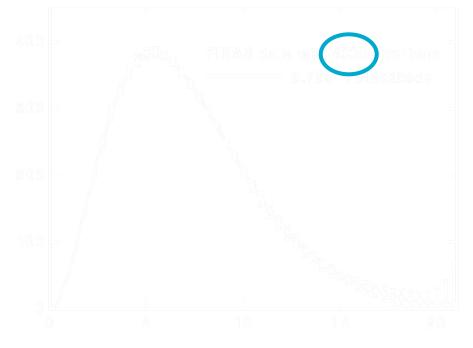






23 Aug 2023

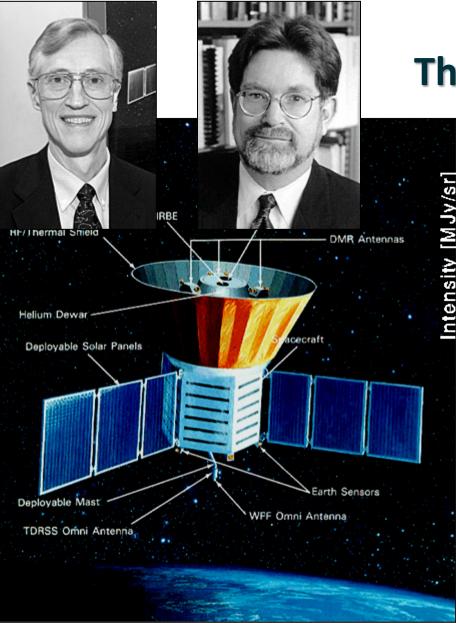


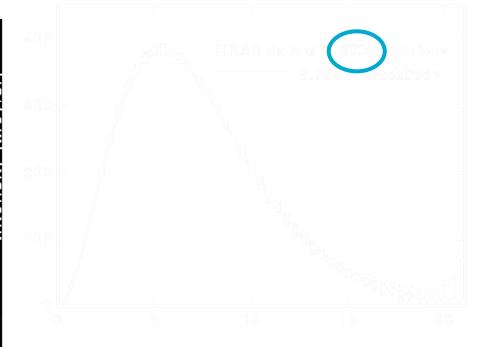


2006 Nobel Prize
 > John Mather – spectrum



History of radio astronomy, URSI GASS: R D Ekers





- 2006 Nobel Prize
 - John Mather spectrum
 - George Smoot anisotropy
 - A rare case of predictions confirmed by observation.



Pulsar discovery: 1967



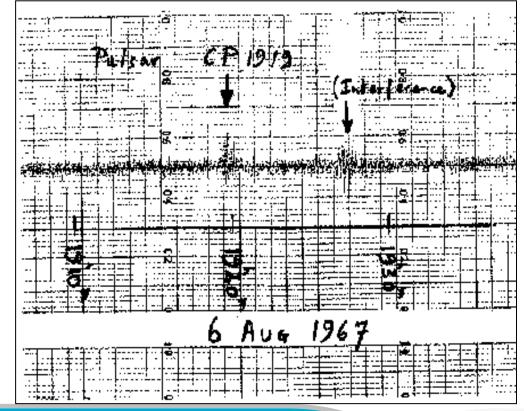
- Jocelyn Bell & Tony Hewish
- Cambridge 1967
- Telescope built for IPS survey
- Short time constant for IPS



Pulsar discovery: 1967



- Jocelyn Bell & Tony Hewish
- Cambridge 1967
- Telescope built for IPS survey
- Short time constant for IPS





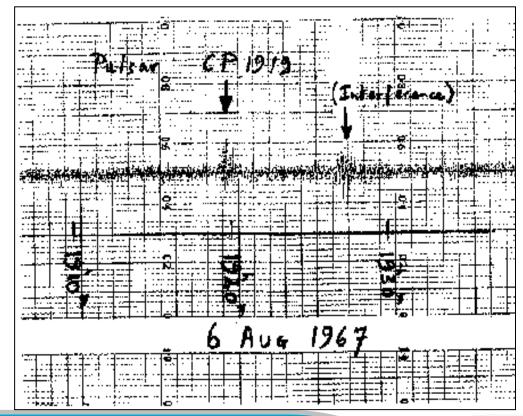
Pulsar discovery: 1967



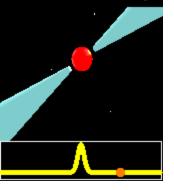
- Jocelyn Bell & Tony Hewish
- Cambridge 1967
- Telescope built for IPS survey
- Short time constant for IPS

• 1974 Nobel Prize to her supervisor Tony Hewish

"A decisive role in the discovery"







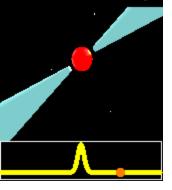
The sound of a Pulsar



History of radio astronomy, URSI GASS: R D Ekers

23 Aug 2023

MPIfR-Bonn Pulsar Group



The sound of a Pulsar

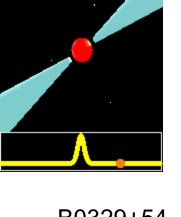
B0329+54



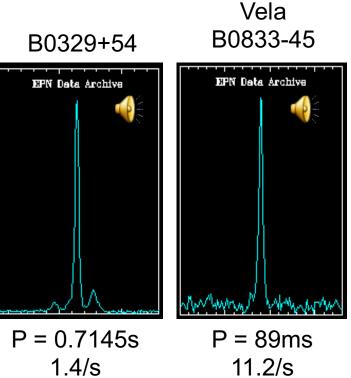
P = 0.7145s 1.4/s



MPIfR-Bonn Pulsar Group



The sound of a Pulsar

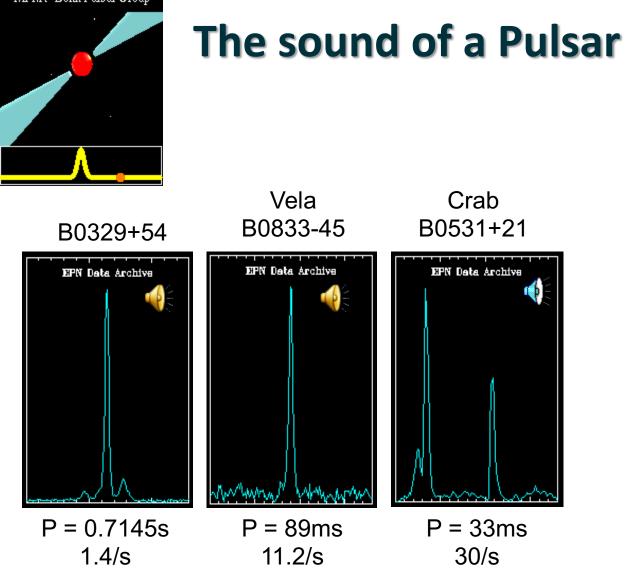




History of radio astronomy, URSI GASS: R D Ekers

23 Aug 2023

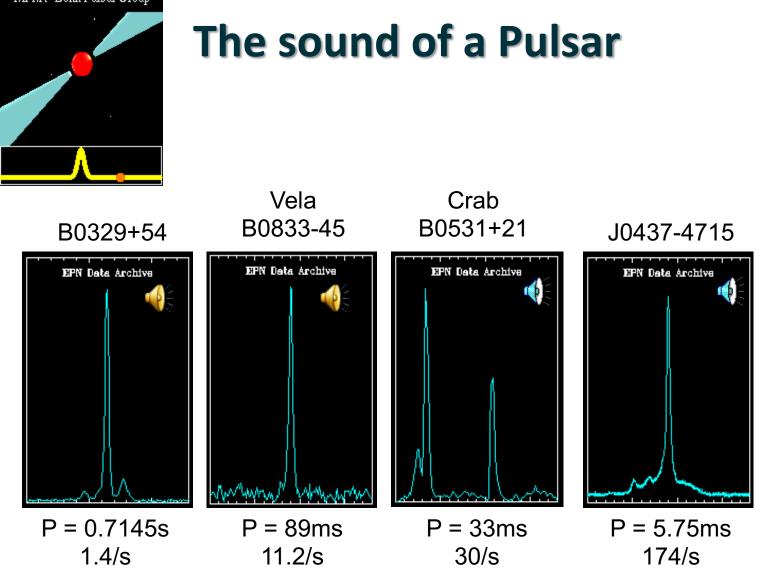
MPIfR-Bonn Pulsar Group





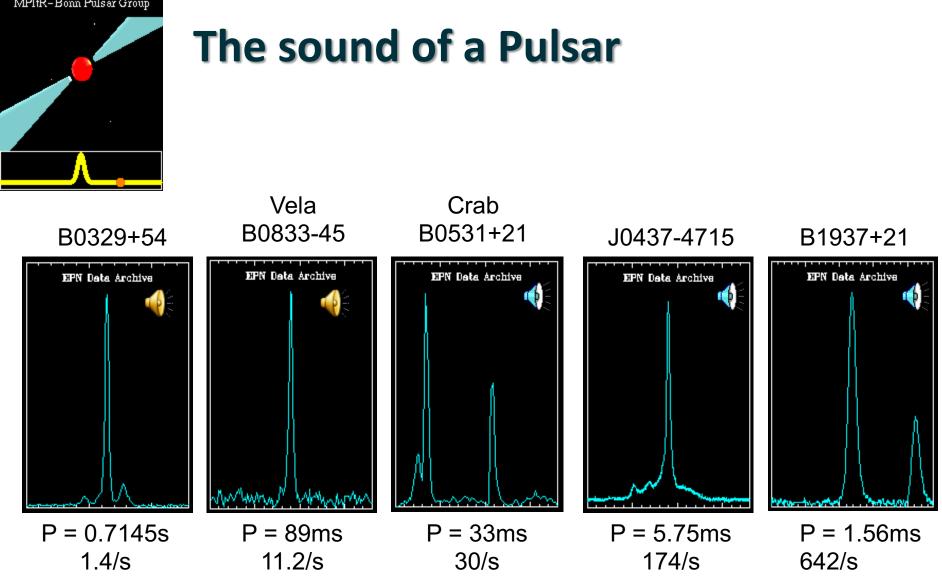
History of radio astronomy, URSI GASS: R D Ekers

MPIfR-Bonn Pulsar Group





MPIfR-Bonn Pulsar Group





History of radio astronomy, URSI GASS: R D Ekers

23 Aug 2023





1993 Noble prize "Gravitational Radiation"



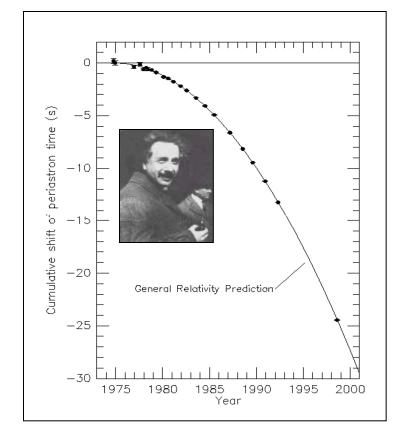
History of radio astronomy, URSI GASS: R D Ekers

23 Aug 2023



1993 Noble prize "Gravitational Radiation"

- 1993 Noble prize to Joe Taylor and Russell Hulse
 - Often considered a classic example of the scientific method
 - But the Nobel citation is for the serendipitous discovery of a binary pulsar
 - Verification of Einstein's prediction of gravitational radiation was not included in the citation!







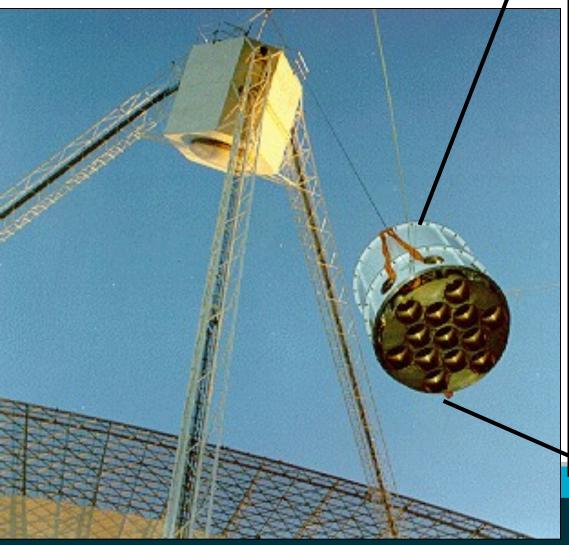
Parkes Multibeam Receiver



- 21 Jan 1997
- Installing the Parkes 21cm Multibeam Receiver
- 13 beams
 - Same as having 13 64m telescopes for surveys!
- HI survey
- Pulsar survey



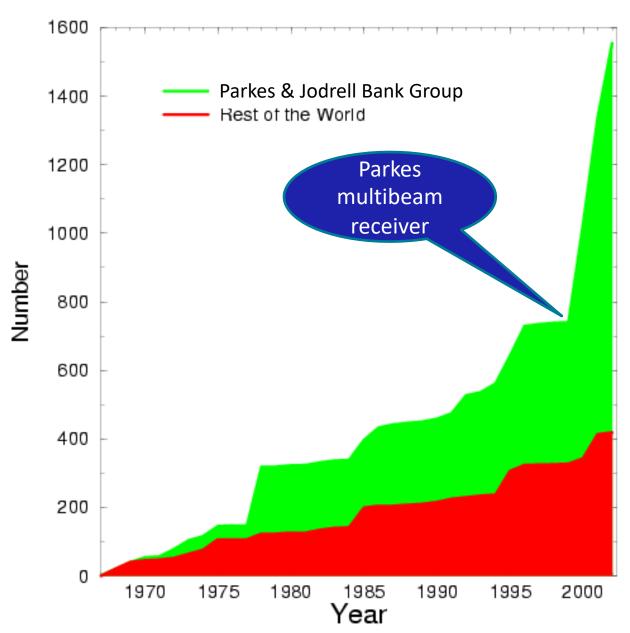
Parkes Multibeam Receiver

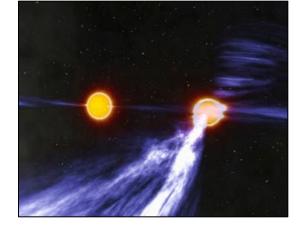




Pulsar discovery rate

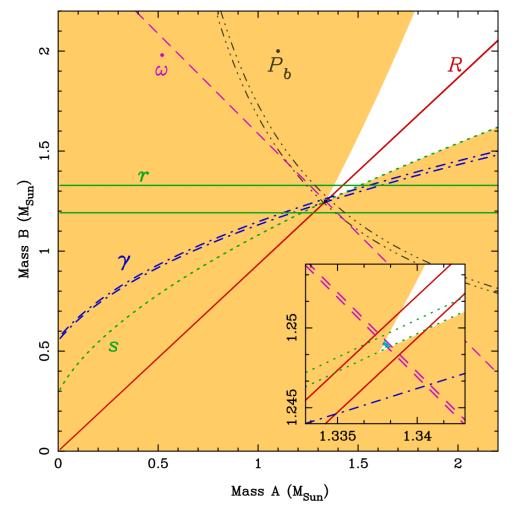
- Dramatic impact of a receiver development
- Using all the focal plane of a big dish





- Parkes Pulsar Survey
- Two neutron stars in 2.4 hour orbit
- General Relativity tested to 0.05%
 - 6 General Relativity parameters tested
 - Kramer et al (2006)

Double Pulsar Tests General Relativity





23 Aug 2023

The impact of Field of View

Lorimer burst and the Parkes multibeam

➢ Discovered 2007

Parkes multibeam - 13 beams = 13 x FoV

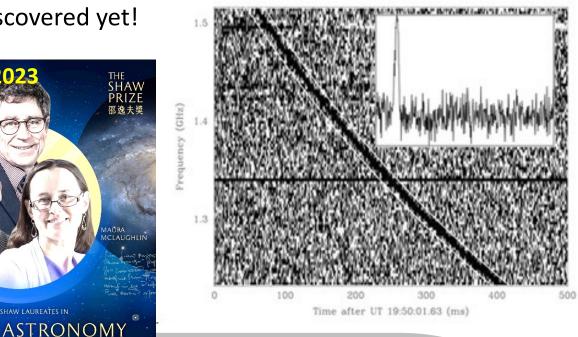
A single beam may have never discovered the Lorimer burst, even though the s/n was huge

➤ Would any FRBs be discovered yet!

Shaw prize 2023

MATTHEW





Bailes

Lorimer

McLaughlan



The impact of Field of View

Lorimer burst and the Parkes multibeam

Discovered 2007

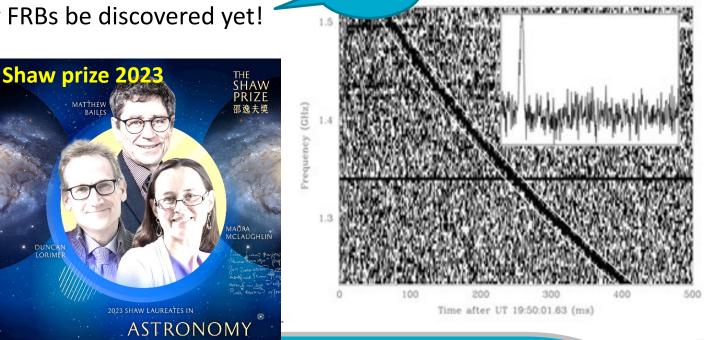
Parkes multibeam - 13 beams = 13 x FoV

 \triangleright A single beam may have never discovered the Lorimer burst, even though the s/n was No!

Would any FRBs be discovered yet!

MATTHEW





Bailes

Lorimer

McLaughlan



CHIME

Line feeds and cylindrical reflectors

➢ Most successful FRB survey telescope



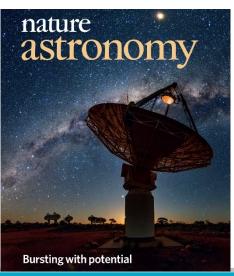


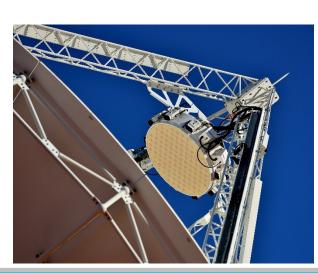
CHIME

- Line feeds and cylindrical reflectors
- ➤ Most successful FRB survey telescope

ASKAP

- ➤ aperture and focal plan array technology
- ➤ 36 12m antennas
- ➤ 36 beams in each antenna
 - $\rightarrow 30 \text{ deg}^2 \text{ per antenna}$









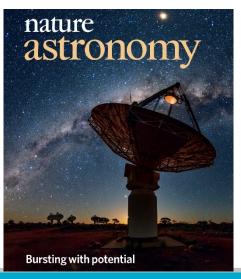


CHIME

- Line feeds and cylindrical reflectors
- ➢ Most successful FRB survey telescope

ASKAP

- ➤ aperture and focal plan array technology
- ≻36 12m antennas
- ➤ 36 beams in each antenna
 - $\rightarrow 30 \text{ deg}^2 \text{ per antenna}$





See ASKAP in Virtual Reality

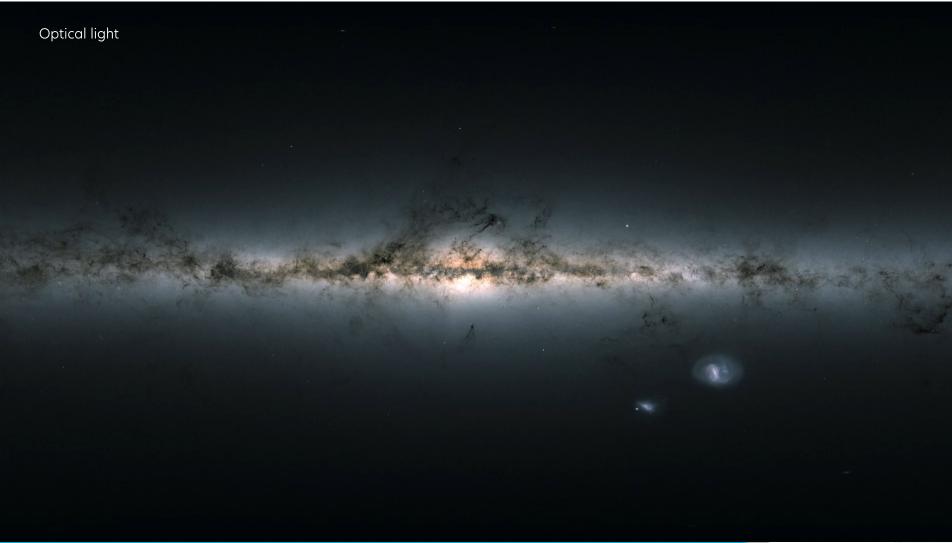






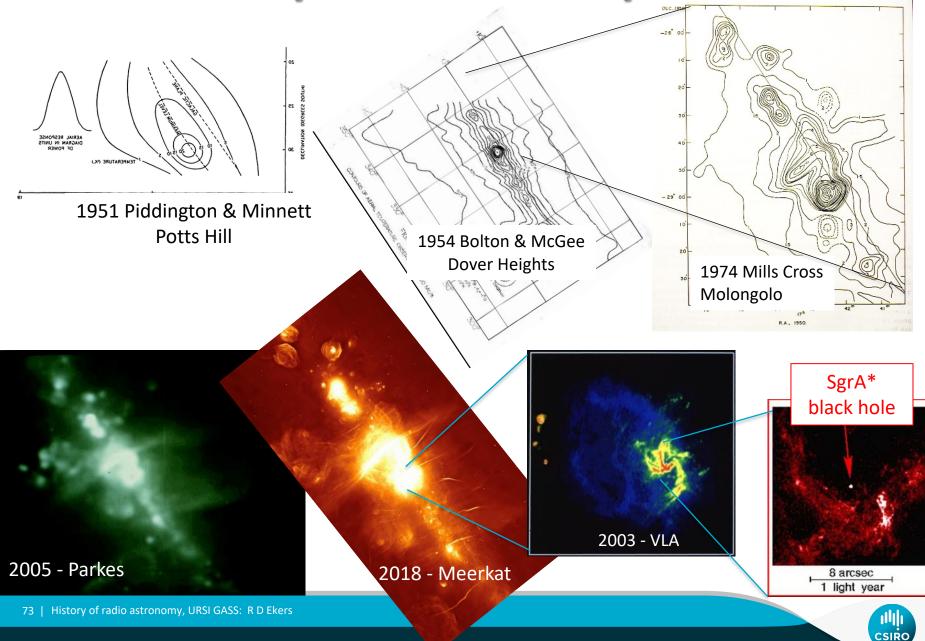
The Event Horizon Telescope team Zoom into the Galactic Centre

video created by Radboud University, Nijmegen

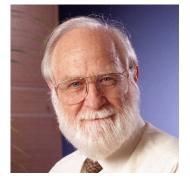


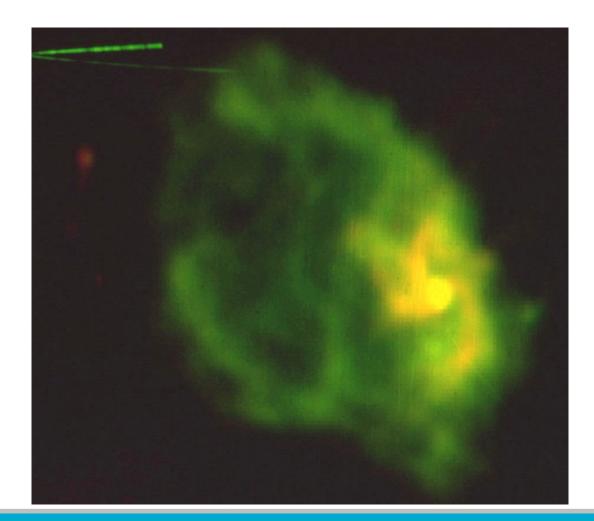


A brief history of the discovery of SgrA*



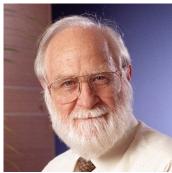
1981 VLA observes our Galactic Centre my greatest discovery moment

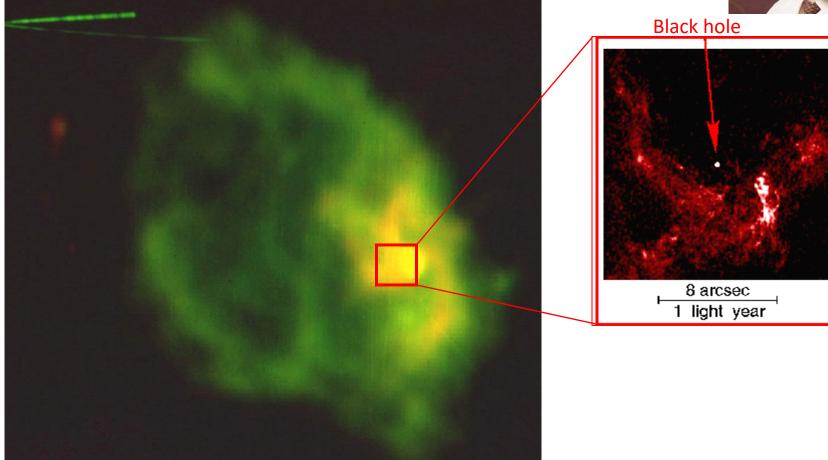






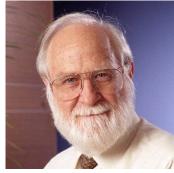
1981 VLA observes our Galactic Centre my greatest discovery moment

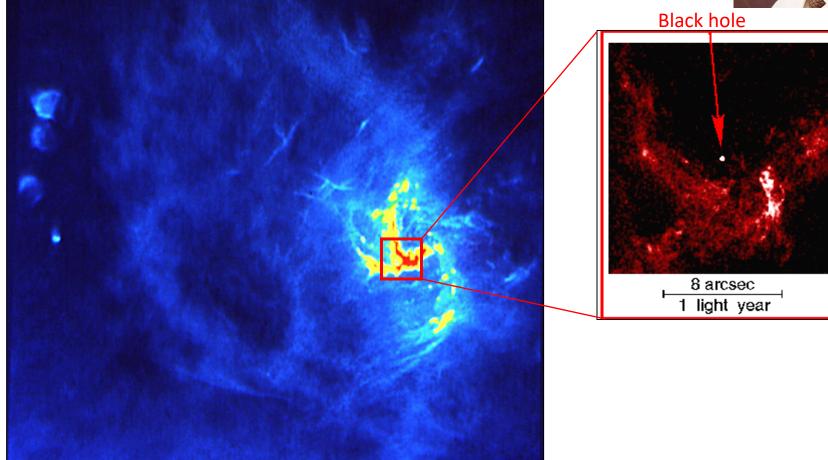






2003 VLA observes Sgr A my greatest discovery moment







MeerKat galactic centre - spectral index

SARAO, Heywood et al. (2022) / J. C. Muñoz-Mateos

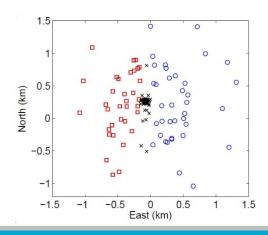
	Spe	stral index	
-1.8	-1.0	0	1.0



76 | History of radio astronomy, URSI GASS: R D Ekers

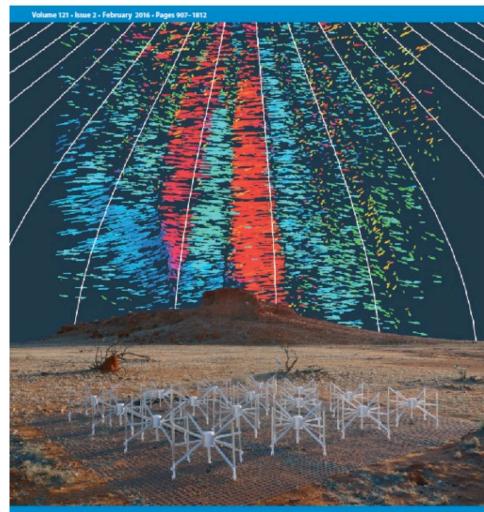
MWA and Ionosphere

- Shyeh Tjing Loi (Cleo) and Tara Murphey
- Transient search leading to the discovery of ionospheric ducts
- Cleo split the MWA to make a parallax measurement



JOURNAL OF GEOPHYSICAL RESEARCH Space Physics

AN AGU JOURNAL



CAGU PUBLICATIONS

WILEY

JGR



Space weather and Interplanetary Scintillation

- URSI General Lecture 3 by Craig Rodger
 - Space weather disturbances in electrical power networks: Space weather features in MWA IPS data
 - ≻John Morgan, Angie Waszewski
 - ➢ New generation Radio Telescopes with very wide FoV essential

